

One Hundred Years of General Relativity

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UCD School of Physics 05/05/2016

100 years of general relativity



First covariant formulation of the field equations (PAS Nov 1915)

First formulation of the full theory (Ann. Phys. May 1916)

844 Sitzung der physikalisch-mathematischen Klasse vom 25. November 1915

Die Feldgleichungen der Gravitation.

VON A. EINSTEIN.

In zwei vor kurzem erschienenen Mitteilungen^[1] habe ich gezeigt, wie man zu Feldgleichungen der Gravitation gelangen kann, die dem Postulat allgemeiner Relativität entsprechen, d. h. die in ihrer allgemeinen Fassung beliebigen Substitutionen der Raumzeitvariablen gegenüber kovariant sind.

Der Entwicklungsgang war dabei folgender. Zunächst fand ich Gleichungen, welche die NEWTONSCHE Theorie als Näherung enthalten und beliebigen Substitutionen von der Determinante 1 gegenüber kovariant waren. Hierauf fand ich, daß diesen Gleichungen allgemein kovariante entsprechen, falls der Skalar des Energietensors der »Materie« verschwindet. Das Koordinatensystem war dann nach der einfachen Regel zu spezialisieren, daß $\sqrt{-g}$ zu 1 gemacht wird, wodurch die Gleichungen der Theorie eine eminentere Vereinfachung erfahren. Dabei mußte aber, wie erwähnt, die Hypothese eingeführt werden,

1916.

№. 7.

ANNALEN DER PHYSIK

VIERTE FOLGE. BAND 49.

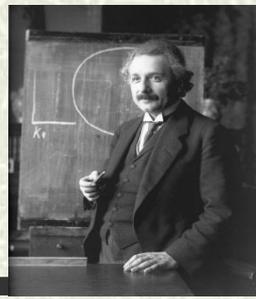
1. Die Grundlage der allgemeinen Relativitätstheorie; von A. Einstein.

Die im nachfolgenden dargelegte Theorie bildet die denkbar weitgehendste Verallgemeinerung der heute allgemein als „Relativitätstheorie“ bezeichneten Theorie; die letztere nenne ich im folgenden zur Unterscheidung von der ersteren „spezielle Relativitätstheorie“ und setze sie als bekannt voraus. Die Verallgemeinerung der Relativitätstheorie wurde sehr erleichtert durch die Gestalt, welche der speziellen Relativitätstheorie durch Minkowski gegeben wurde, welcher Mathematiker zuerst die formale Gleichwertigkeit der räumlichen Koordinaten und der Zeitkoordinate klar erkannte und für

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Overview



The road to general relativity

The special theory of relativity (1905)

The general theory of relativity (1907-16)

Three classic tests (1916)

The perihelion of Mercury; the bending of starlight

The gravitational redshift



Cosmological predictions (1916-18)

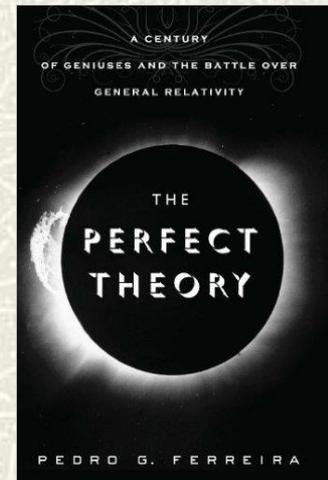
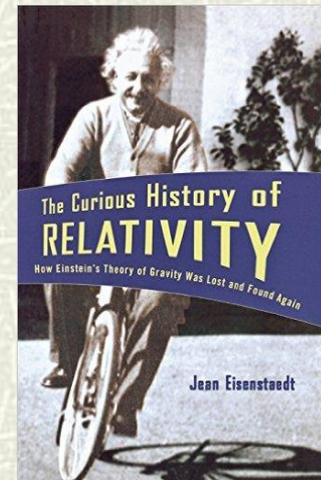
Black holes; gravitational waves

The dynamic universe; the big bang model

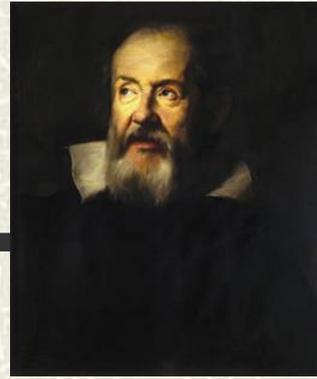
A curious history

Initial success followed by low watermark (1930-60)

The renaissance 1960- astronomy catches up



Relativity



Galileo (1564-1642)

The principle of relativity

Equivalence of a system at rest or moving at constant speed?

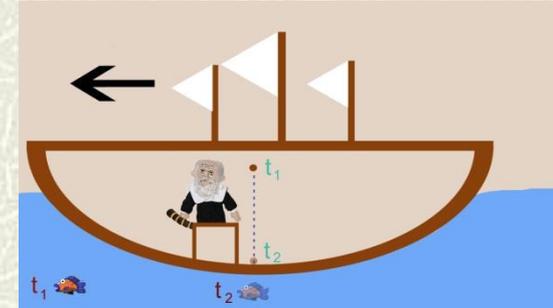
Impossible to distinguish by experiments within the system

Buridan, Oresme, Bruno, Copernicus

Galileo's galleon (1632)

Motion of objects in closed cabin of ship

Independent of motion of ship

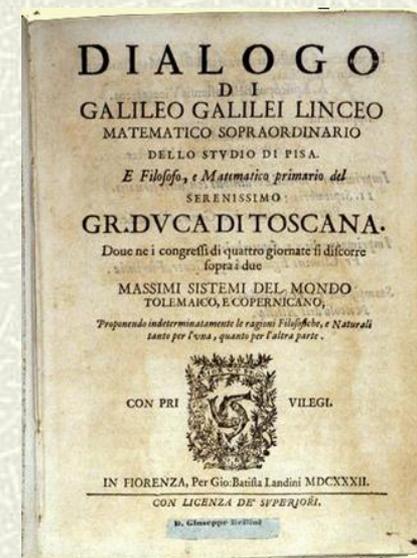


Implication for cosmology

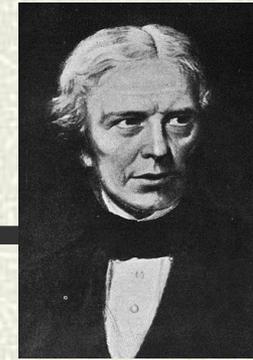
Motion of vessel (earth) undetectable to passengers

Implication for mechanics

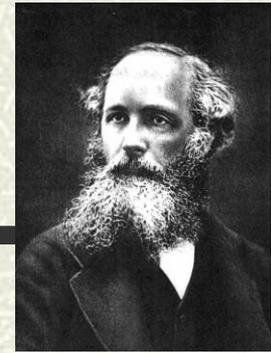
Anticipates Newton's law of inertia



Relativity in the 19th century



Michael Faraday



JC Maxwell

Electromagnetism

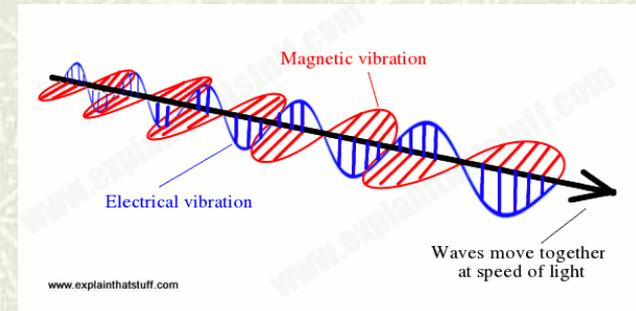
Electricity and magnetism = electromagnetism

Speed of electromagnetic wave = speed of light in vac

Light = an electromagnetic wave

Changing electric and magnetic fields

The electromagnetic spectrum



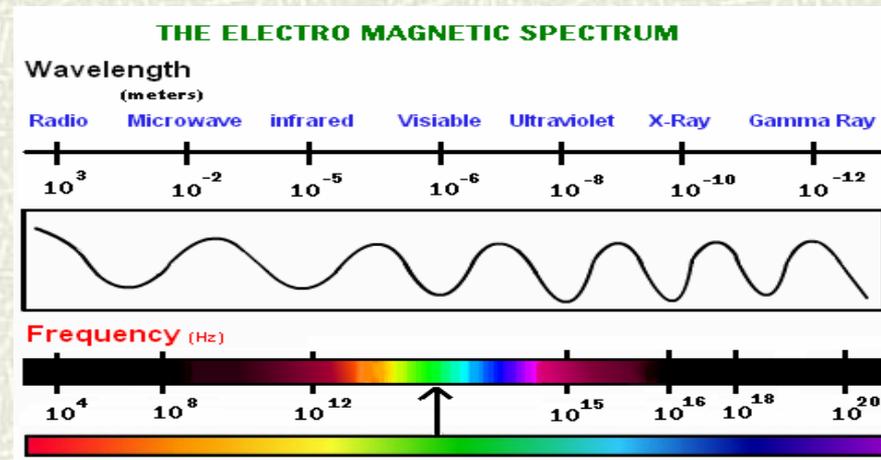
Speed relative to what?

The concept of the ether

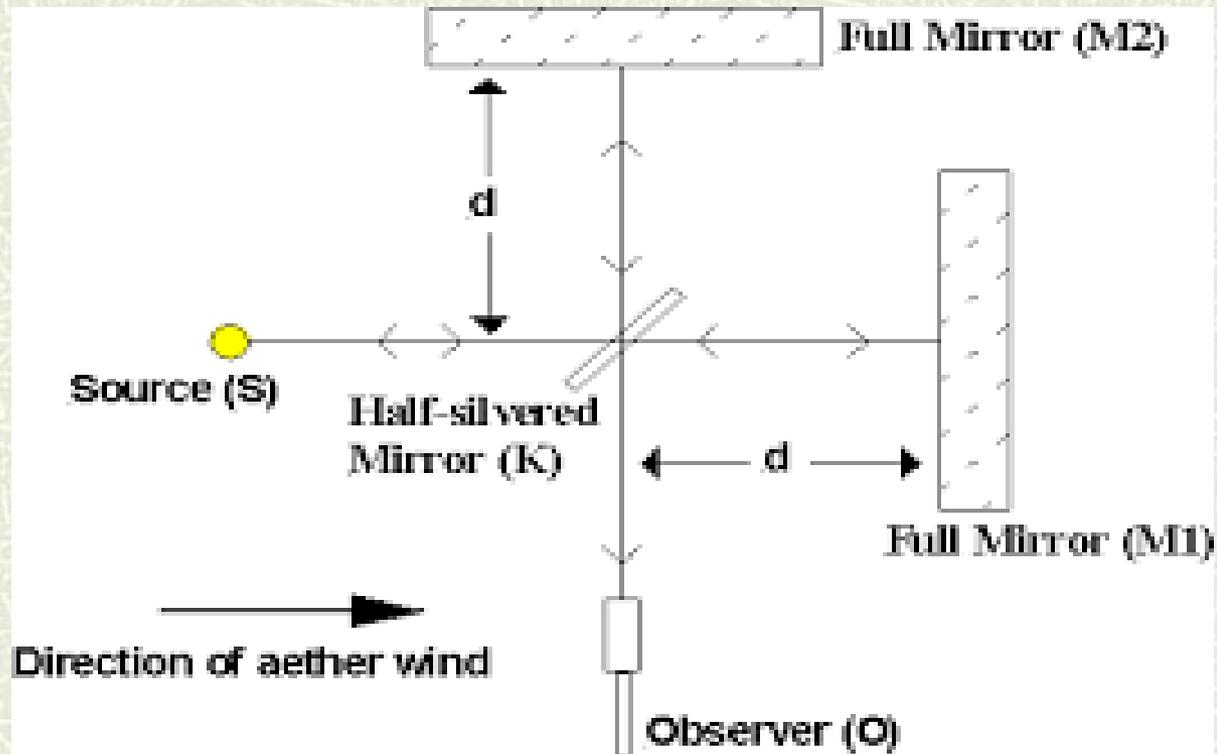
Motion of observer relative to source?

Michelson-Morley experiment

Motion of earth relative to ether?



Michelson-Morley experiment



Expect: rays should arrive at O out of phase

Result: no effect detected

Einstein's special theory of relativity



Two new principles (1905)

Laws of all physics identical for observers in relative uniform motion

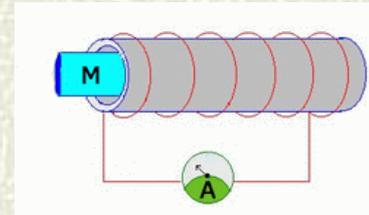
Speed of light in vacuum identical for observers in relative uniform motion

Strange implications

Distance and time not universal

$$v = ds/dt$$

Experienced differently by bodies in relative uniform motion



Predictions for high-speed bodies

Length contraction : time dilation

Mass increase; mass-energy equivalence

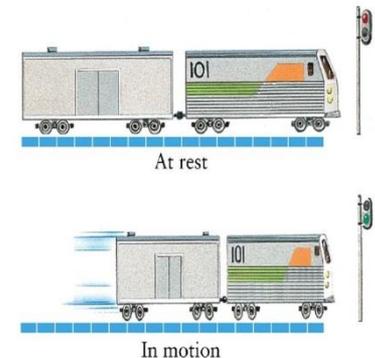
$$E = mc^2$$

Minkowski formulation (1908)

Space-time invariant for observers in inertial frames

$$ds^2 = dx^2 + dy^2 + dz^2 - c^2 dt^2$$

Special Relativity: Length Contraction

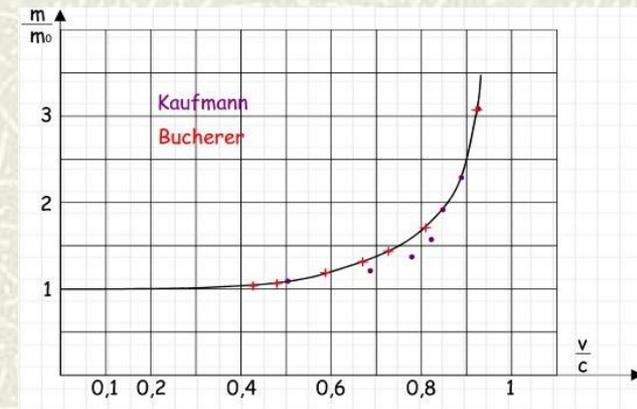


Evidence for special relativity

Mass increase (1901-1912)

The experiments of Kaufmann and Bucherer

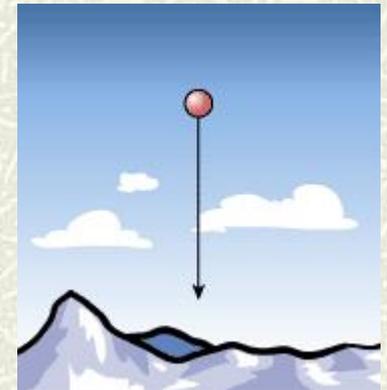
$$m' = \frac{m_0}{\sqrt{1 - v^2/c^2}}$$



Time dilation

The long-lived muon $2 \mu s \rightarrow 22 \mu s$

$$t' = \frac{t_0}{\sqrt{1 - v^2/c^2}}$$



Particle physics experiments

Maximum velocity = c

Mass increase x 1000

Length contraction

Particle creation

$$E = mc^2$$

Special relativity and the media

'Relativity disproven' (Dingle)

Asymmetric controversies



The general theory of relativity (1915,16)

General relativity

Relativity and accelerated motion? (1907)

Relativity and gravity?

The principle of equivalence

Equivalence of gravity and acceleration

Mach's principle

Inertial mass defined relative to matter

"No such thing as empty space"

A long road (1907-1915)

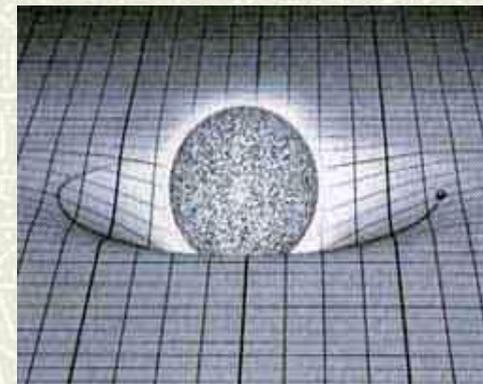
Space-time defined by mass

Gravity = curvature of space-time

Curvilinear geometry, tensor algebra



Falling man doesn't
feel his weight



The field equations of GR (1915)



$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = - \frac{8\pi G}{c^4} T_{\mu\nu}$$

$$G_{\mu\nu} = - \frac{8\pi G}{c^4} T_{\mu\nu}$$

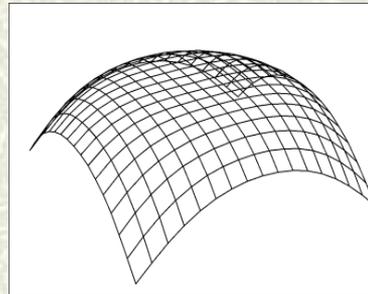


10 non-linear differential equations that relate the spacetime metric to the density and flow of energy and momentum in a region

SR $ds^2 = dx^2 + dy^2 + dz^2 - c^2 dt^2$

$$ds^2 = \sum_{\mu, \nu=1}^4 n_{\mu\nu} dx^\mu dx^\nu$$

$$n_{\mu\nu} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \end{bmatrix}$$



GR $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$

$$ds^2 = \sum_{\mu, \nu=1}^4 g_{\mu\nu} dx^\mu dx^\nu$$

$g_{\mu\nu}$ are functions of the co-ordinates x

Three classic tests for GR (1916)

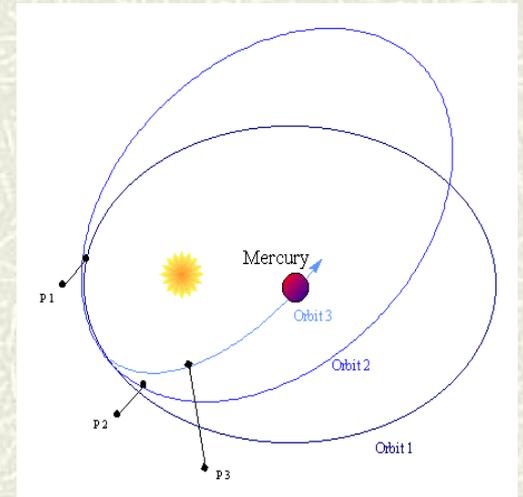
⚡ Different in principle from Newton's gravity

But small deviations in predictions (weak scale)

⚡ The perihelion of Mercury

Well-known anomaly in Mercury's orbit (43" per century)

Correctly predicted by GR (1916)



⚡ The bending of starlight by the sun (1.7")

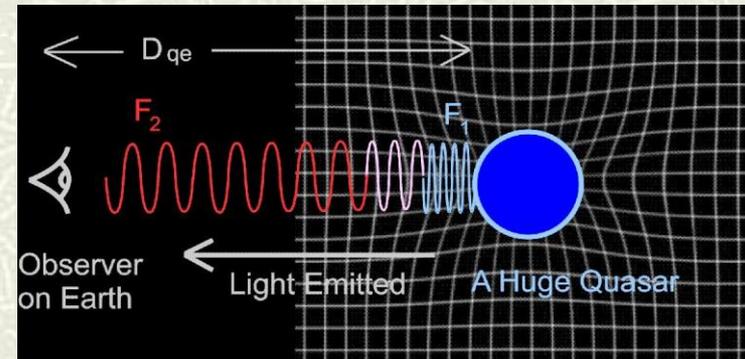
Eclipse expeditions of Eddington and Dyson (1919)

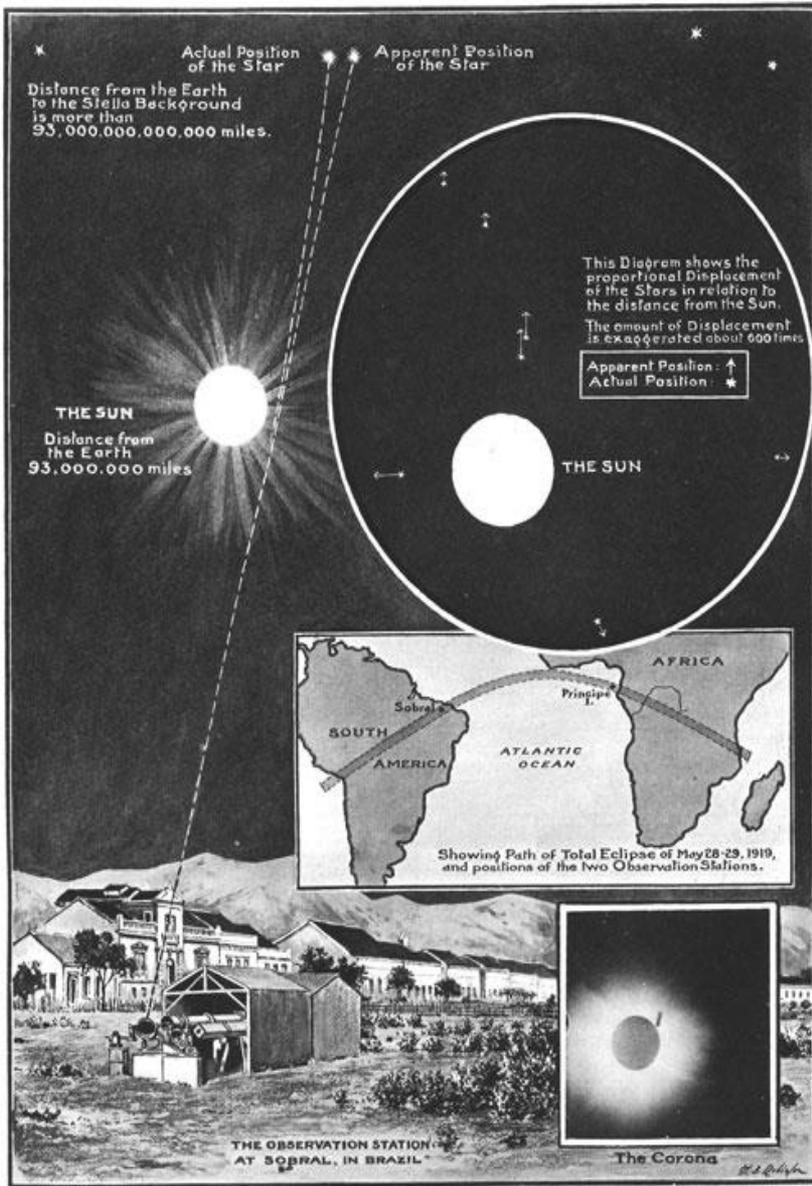
Successful measurement (within error margin)

⚡ Gravitational redshift

Time dilation in strong gravitational field

Light from a massive star redshifted?





Eclipse Results (1919)

Sobral: 1.98" +/- 0.16

Principe: 1.7" +/- 0.4

Einstein famous (1919)

LIGHTS ALL ASKEW IN THE HEAVENS

Men of Science More or Less
Agog Over Results of Eclipse
Observations.

EINSTEIN THEORY TRIUMPHS

Stars Not Where They Seemed
or Were Calculated to be,
but Nobody Need Worry.

A BOOK FOR 12 WISE MEN

No More in All the World Could
Comprehend It, Said Einstein When
His Daring Publishers Accepted It.

Controversy (Collins and Pinch 1970s)
Claim of bias; rebutted by astronomers (RAS)

Three classic tests (1916)

Albert Einstein, The Times (Nov 28th 1919)

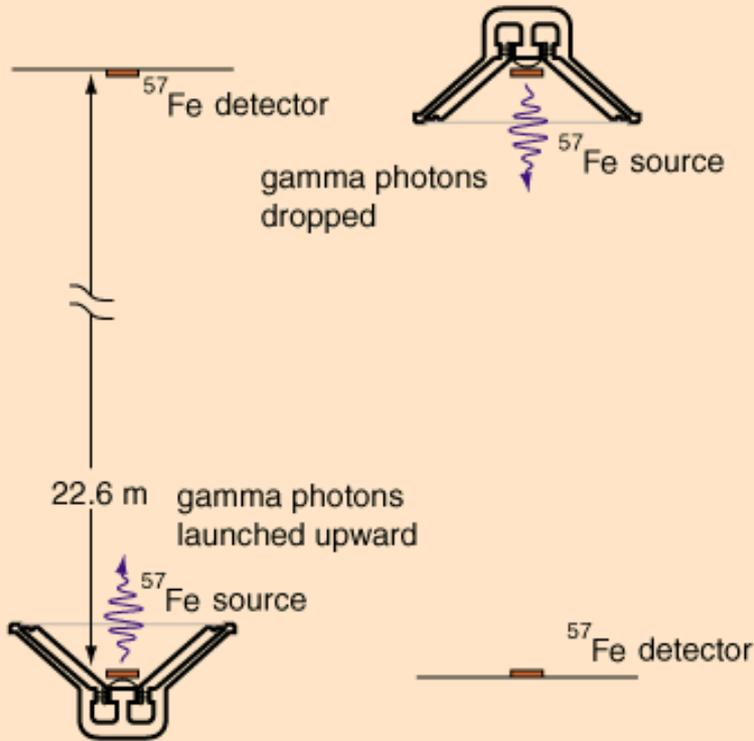
The new theory of gravitation diverges considerably, as regards principles, from Newton's theory. But its practical results agree so nearly with those of Newton's theory that it is difficult to find criteria for distinguishing them which are accessible to experience. Such have been discovered so far:

1. In the revolution of the ellipses of the planetary orbits round the sun (confirmed in the case of Mercury).
2. In the curving of light rays by the action of gravitational fields (confirmed by the English photographs of eclipses).
3. In a displacement of the spectral lines toward the red end of the spectrum in the case of light transmitted to us from stars of considerable magnitude (unconfirmed so far).*

Let no one suppose, however, that the mighty work of Newton can really be superseded by this or any other theory. His great and lucid ideas will retain their unique significance for all time as the foundation of our whole modern conceptual structure in the sphere of natural philosophy.

Gravitational redshift (1959)

Harvard Tower Experiment



In just 22.6 meters, the fractional gravitational red shift given by

$$\nu = \nu_0 \left[1 + \frac{gh}{c^2} \right]$$

is just 4.92×10^{-15} , but the Mössbauer effect with the 14.4 keV gamma ray from iron-57 has a high enough resolution to detect that difference. In the early 60's physicists Pound, Rebka, and Snyder at the Jefferson Physical Laboratory at Harvard measured the shift to within 1% of the predicted shift.

- Sirius B (Adams, 1925)
- Gravity Probe A (1970s)
- Redshifted light from quasars (1990s)

Cosmic prediction I: Black Holes

Schwarzschild (1916)

Exact solution for the field equations

Body of spherical symmetry

Enigma

Solution becomes singular at $r = 0$, $r = 2GM/c^2$

Space closed up around mass?

Rejected

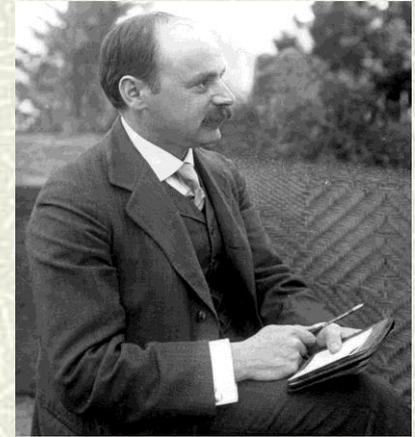
Co-ordinate problem (Eddington)

Prevented by internal pressure (Einstein 1922)

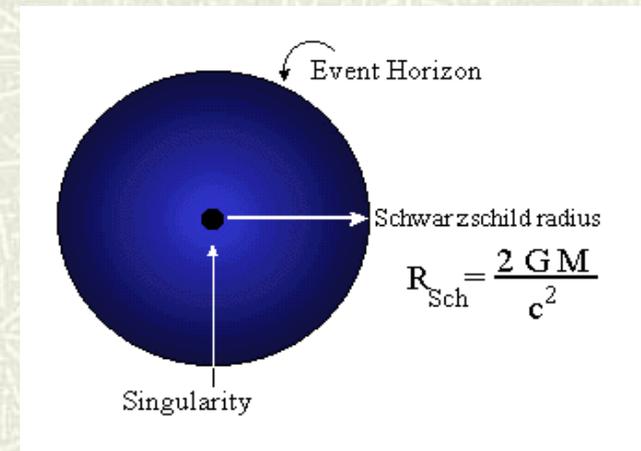
Physical reality?

Collapse of sun? Anderson (UCG)

Collapse of large stellar ensemble : Lodge (Oxford)



Karl Schwarzschild (1873–1916)



The physics of black holes

Chandrasekhar (1931)

The physics of white dwarf stars (quantum degeneracy)

SR: collapse to infinite density for $M > 1.4 M_{\odot}$

Rejected by Eddington, community



Oppenheimer (1939,40)

GR: Continued stellar collapse for $M > 3 M_{\odot}$

Rejected by Einstein (1939)



Wheeler, Thorne, Zeldovitch (1960s)

Numerical solutions of the field equations

Simulation of stellar collapse

Penrose (1965)

No avoiding BH singularity

Black Holes: Observation

Compact astronomical objects (1960s)

Quasars: small, distant sources of incredible energy (1963)

Pulsars: rapidly rotating neutron stars (1967)

X-ray binaries

Cygnus X-1 (1964)

Matter pulled from star into massive companion emits X-rays

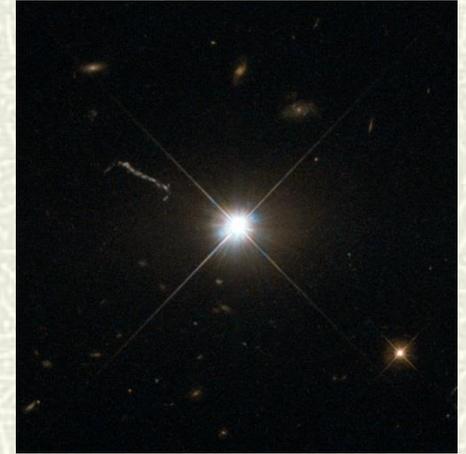
Orbit studies

Supermassive BH at centre of MW? (1990s)

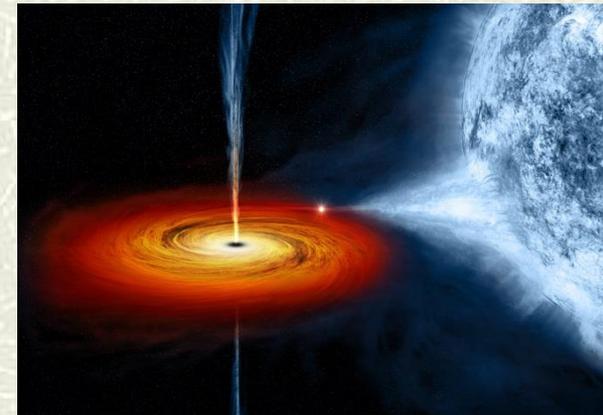
Supermassive BH at centre of many galaxies (2000-)

2015-16

Gravitational waves from binary BH system!



Quasar 3C273



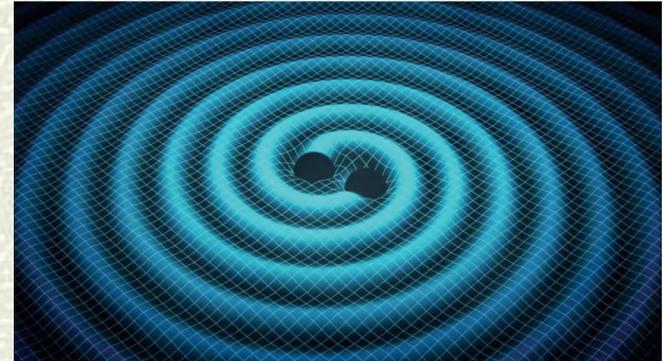
Cygnus X-1 (1964)

Cosmic prediction II: gravitational waves

Einstein (1916, 18)

Linearized wave-like solutions of GFE

Cosmic events cause tiny ripples in spacetime?



Einstein and Rosen (1936, 37)

Cylindrical wave solutions - no gravitational waves (1936)

Corrected with assistance from HP Robertson (1937)

Joseph Weber

Wheeler (1960s)

Numerical wave solutions

Weber bars (1960s)

Reports signal of gravitational waves

Not reproduced, not accepted by community



Gravitational Waves: Observation

Binary pulsar PSR1913+16

Hulse-Taylor (1974)

Decrease in orbital period exactly as predicted

Direct measurement?

Interferometers: 1980-

Interferometers with 4 km arms (LIGO, VIRGO)

Advanced LIGO (2015)

Clear signal (September 2015)

BH merger

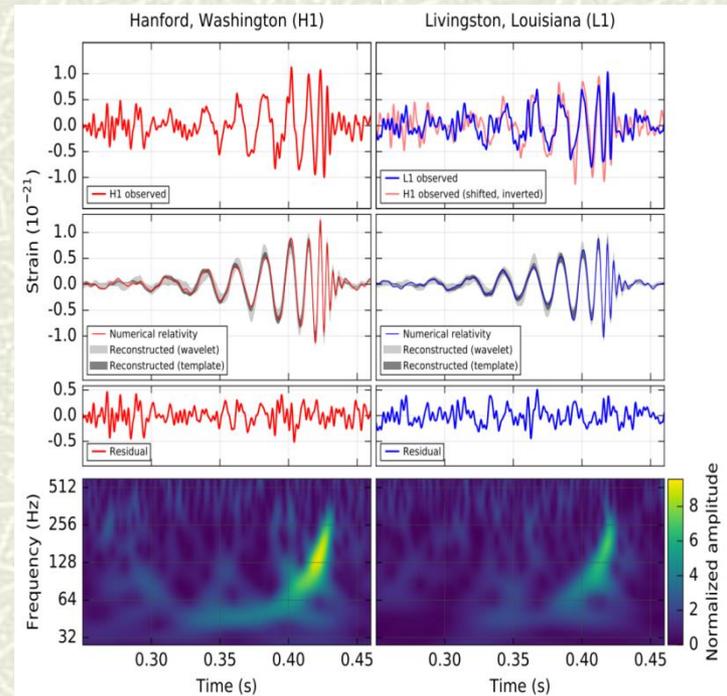
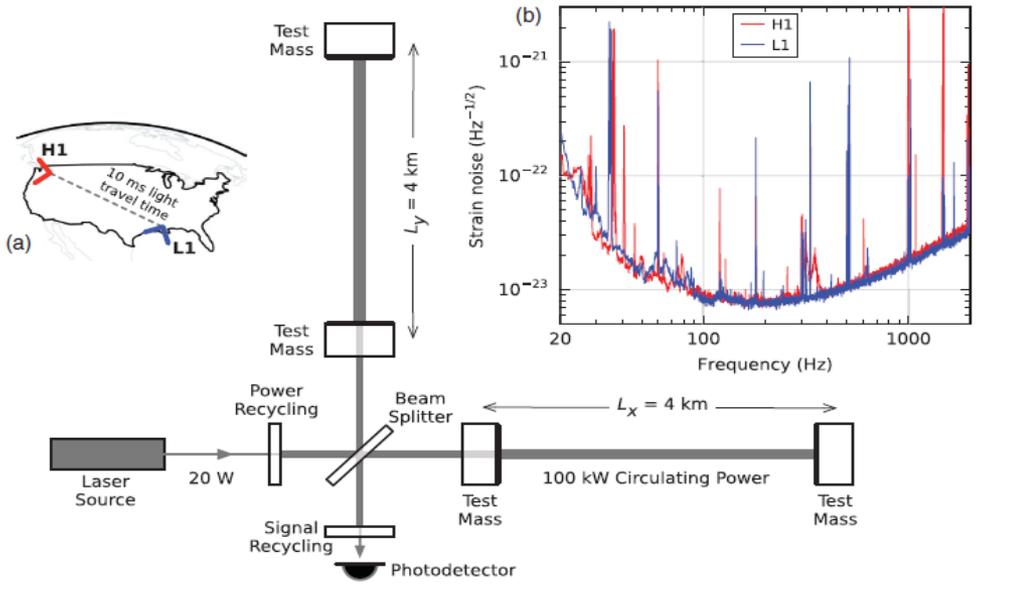
Exact match with merging BHs

$29 M_{\odot}, 36 M_{\odot}$; 1.3 billion LY away



Hulse-Taylor pulsar





1. Shape of waveform
2. Frequency of orbit

BBH !

061102 (2016)

Selected for a Viewpoint in *Physics*
 PHYSICAL REVIEW LETTERS

week
 12 FEBRU



Observation of Gravitational Waves from a Binary Black Hole Merger

B. P. Abbott *et al.**

(LIGO Scientific Collaboration and Virgo Collaboration)

(Received 21 January 2016; published 11 February 2016)

On September 14, 2015 at 09:50:45 UTC the two detectors of the Laser Interferometer Gravitational-Wave Observatory simultaneously observed a transient gravitational-wave signal. The signal sweeps upwards in frequency from 35 to 250 Hz with a peak gravitational-wave strain of 1.0×10^{-21} . It matches the waveform predicted by general relativity for the inspiral and merger of a pair of black holes and the ringdown of the resulting single black hole. The signal was observed with a matched-filter signal-to-noise ratio of 24 and a false alarm rate estimated to be less than 1 event per 203 000 years, equivalent to a significance greater than 5.1σ . The source lies at a luminosity distance of 410^{+160}_{-180} Mpc corresponding to a redshift $z = 0.09^{+0.03}_{-0.04}$. In the source frame, the initial black hole masses are $36^{+4}_{-4} M_{\odot}$ and $29^{+4}_{-4} M_{\odot}$, and the final black hole mass is $62^{+4}_{-4} M_{\odot}$, with $3.0^{+0.5}_{-0.5} M_{\odot} c^2$ radiated in gravitational waves. All uncertainties define 90% credible intervals. These observations demonstrate the existence of binary stellar-mass black hole systems. This is the first direct detection of gravitational waves and the first observation of a binary black hole merger.

Cosmic prediction III: The expanding universe

Einstein: apply GR to the Universe (1917)

Ultimate test for new theory of gravitation

Assumptions

Static universe (small velocities of the stars)

Mach's principle (metric tensor to vanish at infinity)

Isotropy and homogeneity (simplicity)

Boundary problem

Assume cosmos of closed curvature

But...no consistent solution

New term in field equations!

Cosmic constant - anti-gravity term

Radius and density defined by λ

$$G_{\mu\nu} = -\kappa T_{\mu\nu}$$

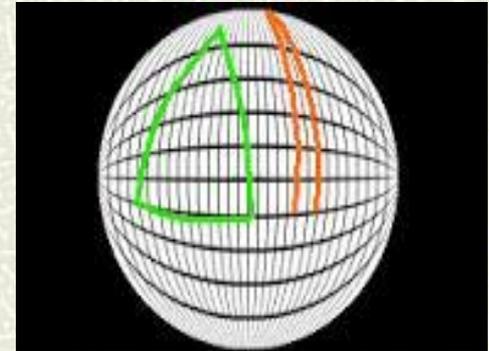
$$G_{\mu\nu} + \lambda g_{\mu\nu} = -\kappa T_{\mu\nu}$$

Doc. 43

Cosmological Considerations in the General Theory of Relativity

This translation by W. Perrett and G. B. Jeffery is reprinted from H. A. Lorentz et al., *The Principle of Relativity* (Dover, 1952), pp. 175–188.

It is well known that Poisson's equation $\nabla^2\phi = 4\pi K\rho$ (1) in combination with the equations of motion of a material point is not as yet a perfect substitute for Newton's theory of action at a distance. There is still to be taken into account the condition that at spatial infinity the potential ϕ tends



$$\lambda = \frac{\kappa\rho}{2} = \frac{1}{R^2}$$

Friedman's universe



Alexander Friedman
(1888 -1925)

Allow time-varying solutions (1922)

Assume homogeneity, isotropy, positive curvature

Two independent differential equations from GFE

$$G_{\mu\nu} + \lambda g_{\mu\nu} = -\kappa T_{\mu\nu}$$

$$\frac{3R'^2}{R^2} + \frac{3c^2}{R^2} - \lambda = \kappa c^2 \rho,$$

Evolving universes

Density of matter varies over time

$$\frac{R'^2}{R^2} + \frac{2RR''}{R^2} + \frac{c^2}{R^2} - \lambda = 0.$$

Negative spatial curvature (1924)

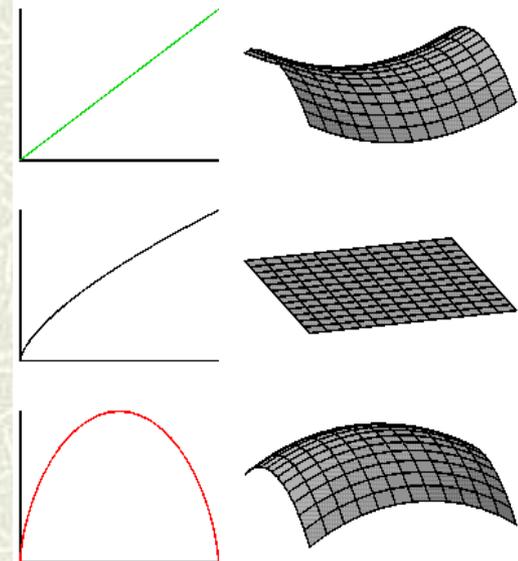
Cosmic evolution, geometry depends on matter

Overlooked by community

Considered 'suspicious' by Einstein

Mathematical correction, later retracted

~~"To this a physical reality can hardly be ascribed"~~



Lemaître's universe (1927)



Expanding model of the cosmos from GR

Similar but not identical to Friedman 1922

$$3 \frac{R'^2}{R^2} + \frac{3}{R^2} = \lambda + \kappa \rho$$

Redshifts of galaxies = expansion of space?

Redshifts from Slipher, distances from Hubble

$$2 \frac{R''}{R} + \frac{R'^2}{R^2} + \frac{1}{R^2} = \lambda - \kappa \rho$$

Fr Georges Lemaître

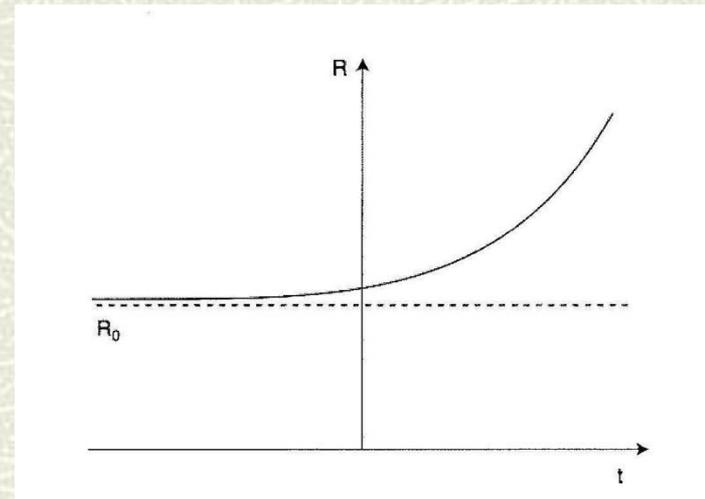
Ignored by community

Belgian journal (in French)

Rejected by Einstein: "Vôtre physique est abominable"

Lemaître's recollection (1958)

"Einstein not up-to-date with astronomy"



The watershed: Hubble's law (1929)



The redshifts of the spiral nebulae

Doppler shifts? (Slipher 1915, 1917)

The distances to the nebulae

Far beyond Milky Way (Hubble 1925)

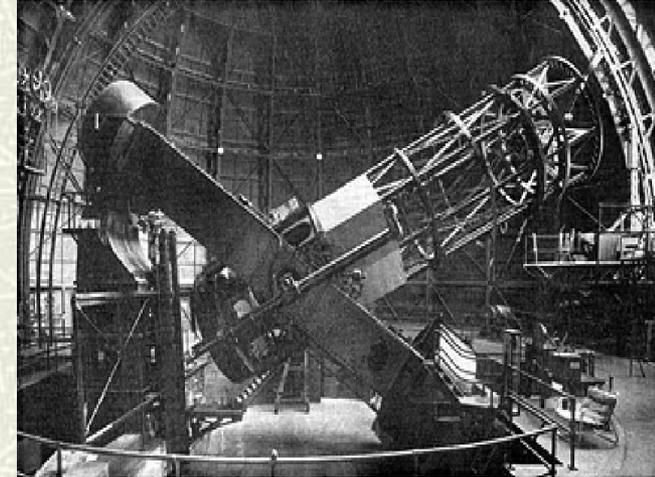
A redshift/distance relation?

Linear relation (Hubble, 1929)

$$H = 500 \text{ kms}^{-1} \text{ Mpc}^{-1}$$

Landmark result in astronomy

Furthest galaxies receding the fastest



Velocity-Distance Relation among Extra-Galactic Nebulae.

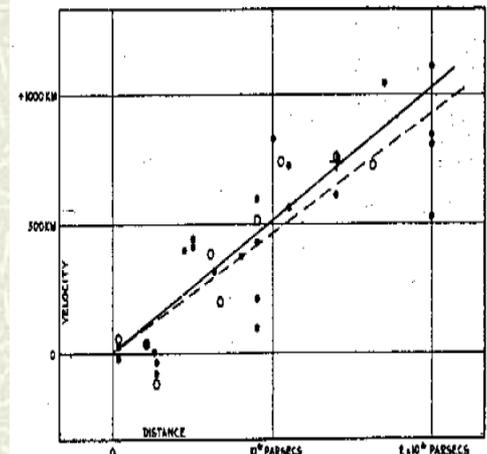


FIGURE 1

What is causing the effect?

The expanding universe (1930)

- **RAS meeting (May 1930)**

*If redshifts are velocities, and if effect is non-local
Hubble's law = expansion of space? (Edd., de Sitter)*

- **Dynamic model required**

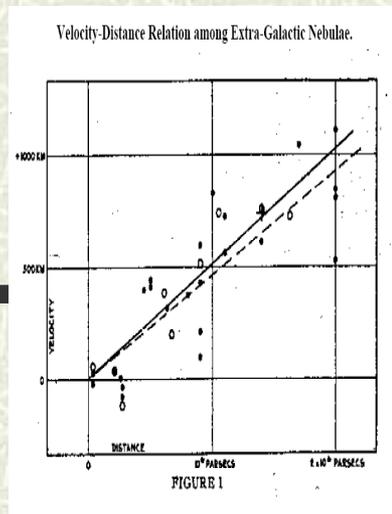
Static model unstable

- **Lemaître model adopted**

*1927 expanding model republished in English (1931)
Observational section omitted (rightly)*

- **Lemaître-Friedman cosmology accepted**

Time-varying radius, decreasing density of matter



Models of the expanding universe (1930 -)

- **Tolman (1930, 31)**

Expansion caused by annihilation of matter ?

- **Eddington (1930, 31)**

*On the instability of the Einstein universe
Expansion caused by condensation?*

- **de Sitter (1930, 31)**

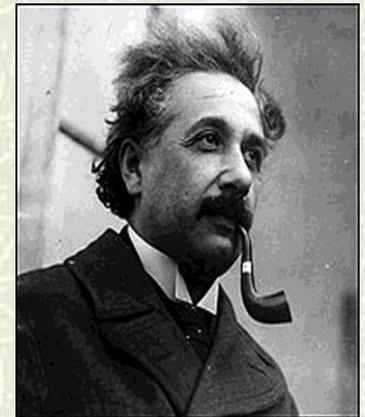
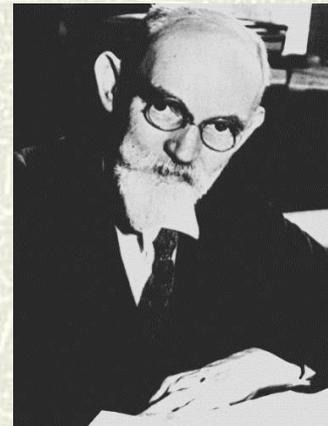
Variety of expanding models

- **Heckmann (1931,32)**

Spatial curvature (not translated)

- **Einstein (1931, 32)**

*Friedman-Einstein model $\lambda = 0, k = 1$
Einstein-de Sitter model $\lambda = 0, k = 0$*



*If redshifts represent expansion...
If effect is global...*

Cosmic prediction IV: the big bang



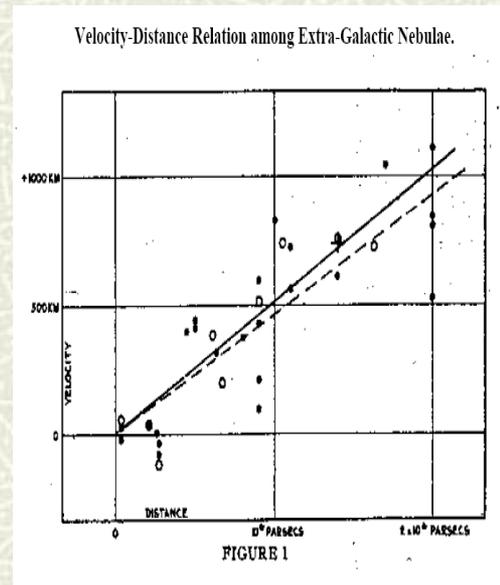
- # **Lemaître 1931**: expanding U smaller in the past
- # Extrapolate to very early epochs
- # Extremely dense, extremely hot
- # Expanding and cooling ever since
- # 'Fireworks beginning' at $R = 0$?

Fr Georges Lemaître

Rejected by community (1930-60)

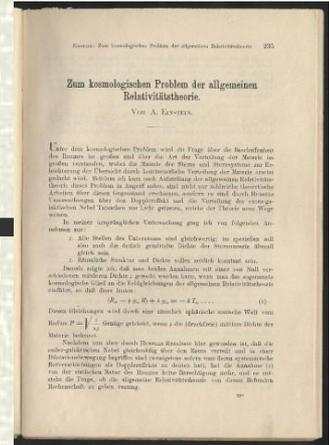
Simplified models

Timescale problem



Later called *'The big bang'*

Einstein's 1931 model



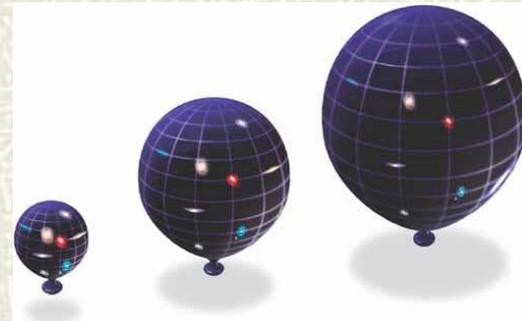
Einstein's first dynamic model of the cosmos

Often cited, rarely read (not translated)

Adopts Friedman 1922 model

Time-varying, closed universe: $k = 1$

Cosmic constant redundant: set $\lambda = 0$



Use Hubble to extract parameters

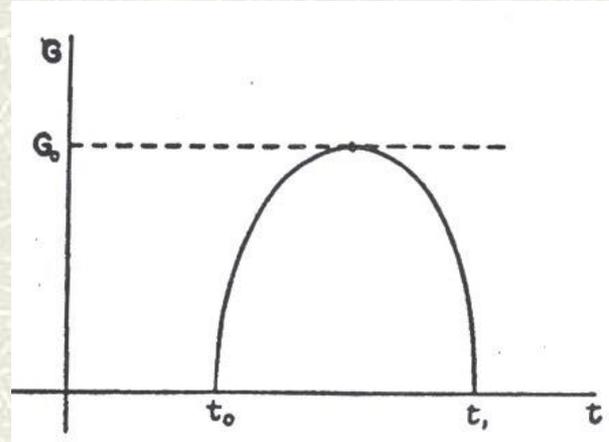
Radius $R \sim 10^8$ lyr

Density of matter $\rho \sim 10^{-26}$ g/cm³

Timespan problem

10^{10} yr: conflict with astrophysics

Attributed to simplifying assumptions (homogeneity)



Einstein's 1931 model revisited

First translation into English

O'Raifeartaigh and McCann 2014

Anomalies in calculations of radius and density

$R \sim 10^8 \text{ lyr} :$ *should be 10^9 lyr*

$\rho \sim 10^{-26} \text{ g/cm}^3 :$ *should be 10^{-28} g/cm^3*

$t \sim 10^{10} \text{ yr} :$ *should be 10^9 yr*

Source of error?

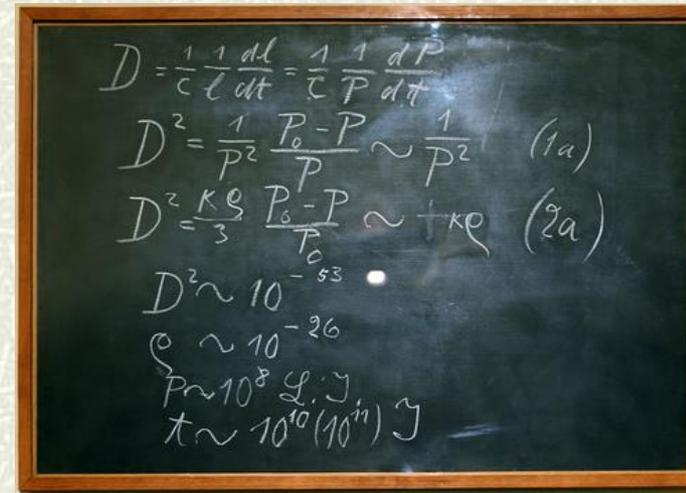
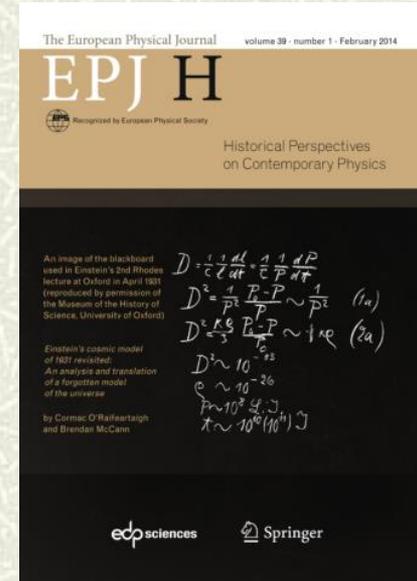
Error in Hubble constant (Oxford blackboard)

$D^2 \sim 10^{-53}$ *instead of 10^{-55} cm^{-2}*

Not a cyclic model

“Model fails at $P = 0$ ”

Contrary to what is often stated



Bonus: Einstein's steady-state model

Unpublished manuscript

Archived as draft of Friedman-Einstein model

Similar title, opening

Something different

Cosmological constant λ

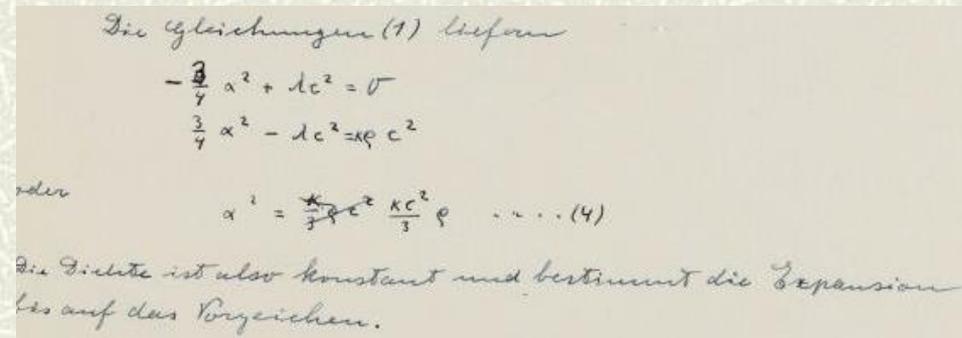
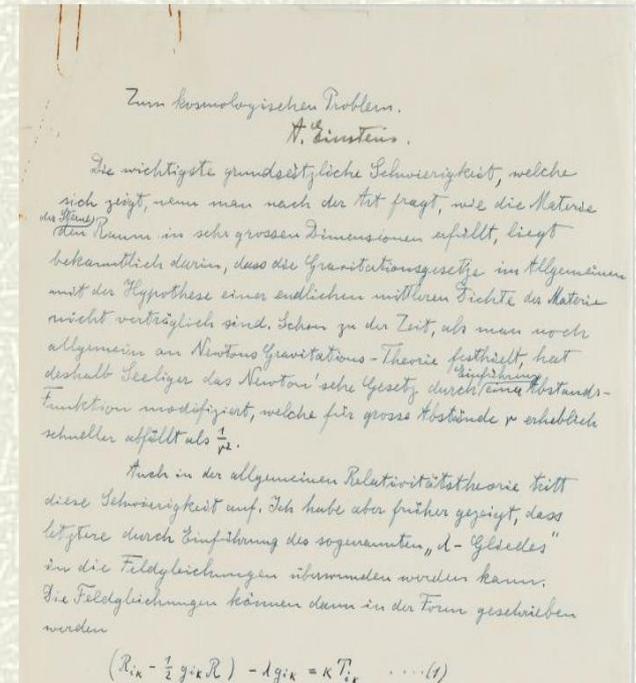
"The density is constant and determines the expansion"

Steady-state model

Continuous formation of matter from vacuum

Anticipates Hoyle's model

Fatal flaw: abandoned



The steady-state universe (1948)

Expanding but unchanging universe

Hoyle, Bondi and Gold (1948)

Disliked extrapolation to early epochs

Perfect cosmological principle?



Bondi, Gold and Hoyle

Requires continuous creation of matter

Very little matter required

No beginning, no age paradox

Replace λ with creation term (Hoyle)

$$G_{\mu\nu} + C_{\mu\nu} = -k T_{\mu\nu}$$

Improved version (1962)

$$G_{\mu\nu} + \lambda g_{\mu\nu} = k T (C_{\mu} + C_{\nu})$$



Hoyle and Narlikar (1962)

Steady-state vs big bang (1950-70)

Optical astronomy (1950s)

Revised distances to the nebulae (Baade, Sandage)

New timescale of expansion

Radio-astronomy (1960s)

Galaxy distributions at different epochs

Cambridge 3C Survey (Ryle)

Nucleosynthesis of light elements

Gamow et al. 1948

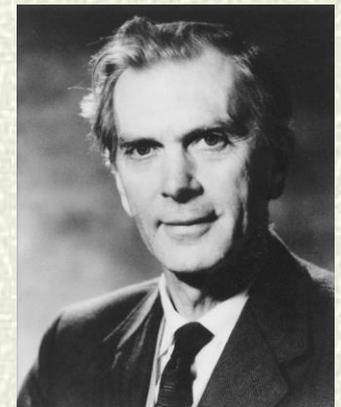
Cosmic microwave background (1965)

Low temperature, low frequency

Remnant of young, hot universe



Allen Sandage



Martin Ryle

NATURE | NEWS   

Einstein's lost theory uncovered

Physicist explored the idea of a steady-state Universe in 1931.

Daive Castelvechi

24 February 2014

Physics » Nature   Email  Print

Einstein's Lost Theory Uncovered

The famous physicist explored the idea of a steady-state universe in 1931

nature

Feb 25, 2014 | By **Daive Castelvechi** and Nature magazine

A manuscript that lay unnoticed by scientists for decades has revealed that **Albert Einstein** once dabbled with an



New Discovery Reveals Einstein Tried To Devise A Steady State Model Of The Universe

www.irishtimes.com/news/science/wit-researchers-discover-lost-einstein-model-of-universe-1.1713487

THE IRISH TIMES **Science** Monday, March 10, 2014

News Sport Business Debate Life & Style Culture Offers

News / Science

2 comments, 2 called-out [+ Comment Now](#) [+ Follow Comments](#)

Almost 20 years before the late Fred Hoyle and his colleagues devised the [Steady State Theory](#), Albert Einstein toyed with a similar idea: that the universe was eternal, expanding outward with a consistent input of spontaneously generating matter.

An Irish physicist came across the paper last year and could hardly believe. According to this week's article in [Nature](#),

model of the universe very different to today's [Big Bang](#) Theory.

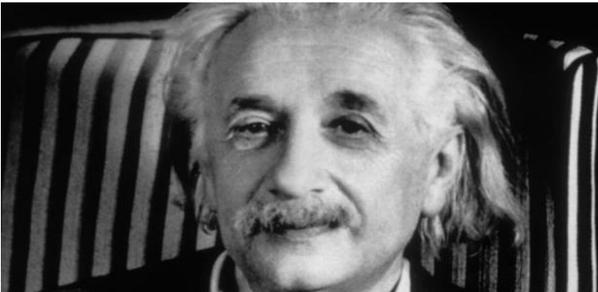
The manuscript, which hadn't been referred to by scientists for decades,



[TheJournal.ie](#) Like You like this.

WIT researchers discover 'lost' Einstein model of universe

Scientists uncovered misfiled papers while searching Jerusalem university's online archive



Latest Ireland »

- 12:26 Quinn confirms Flannery approached hm with Rehab concerns
- 09:07 Man in his twenties stabbed in north Dublin
- 09:05 Family hope public appeal will help daughter beat cancer
- 08:42 Gardaí investigate death of woman in Dublin
- 08:25 Flannery faces call from all parties to attend PAC

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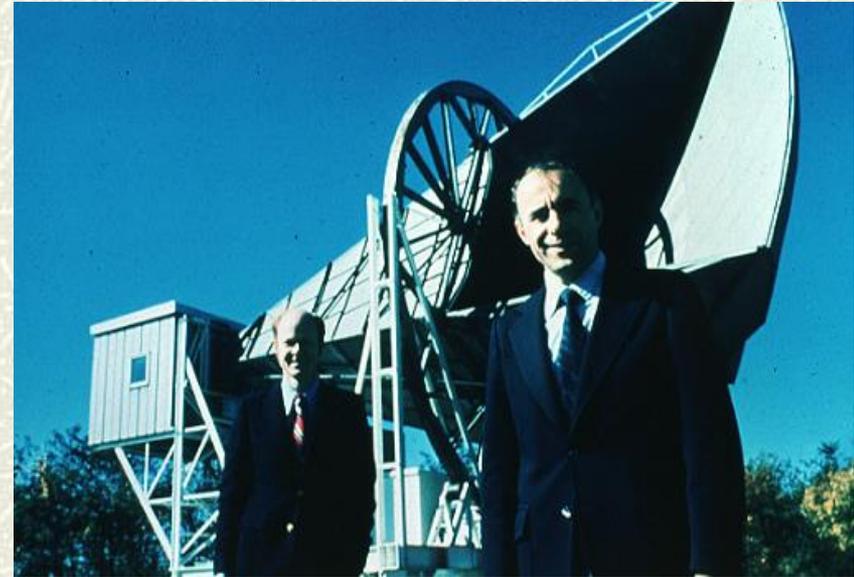
The way back isn't so simple



Cosmic background radiation (exp)

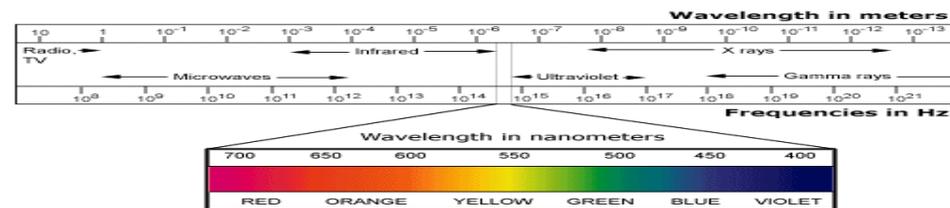
CMB discovered accidentally

- # Universal signal (1965)
- # Low frequency (microwave)
- # Low temperature (3K)



Penzias and Wilson (1965)

Echo of Big Bang!



Modern measurements of the CMB

- Details of background radiation
 - Full spectrum
 - Comparison with theory
 - Perturbations?
-
- *Ground telescopes*
 - *Balloon experiments*
 - *Satellite experiments*



COBE satellite (1992)

Today's cosmological puzzles

Characteristics of background radiation

Isotropy, homogeneity, flatness (1970-80)

The theory of inflation (1981)

Exponential expansion within first second?

Explanation for homogeneity, flatness, galaxy formation

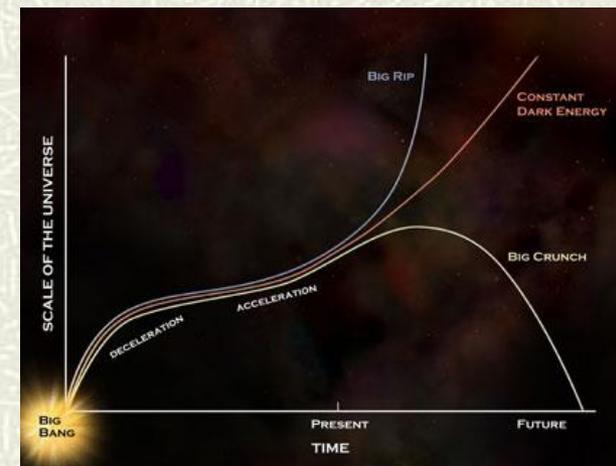
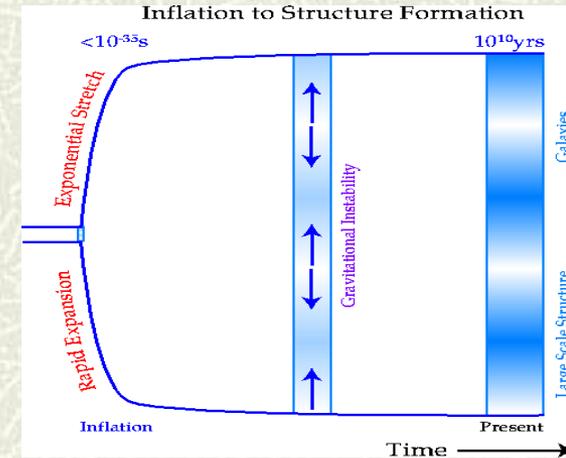
Which model of inflation? Bubble inflation?

Dark energy (1998)

Observation of accelerated expansion

The return of the cosmological constant

U mainly composed of dark energy



$$\mathbf{G}_{\mu\nu} + \lambda g_{\mu\nu} = -\kappa \mathbf{T}_{\mu\nu}$$

Nature of DE unknown

100 years of general relativity

✦ **Published 1915, 1916**

From Swiss patent office to Berlin

✦ **A new theory of gravity**

Gravity = curvature of spacetime

✦ **Classic predictions supported by experiment**

Perihelion of Mercury: bending of light by a star

Gravitational redshift ; GPS

✦ **Cosmological predictions supported by experiment**

Black holes: gravitational waves

The expanding universe: the big bang



Skeptical of extrapolations

Where next for general relativity?

More general theory

Unified field theory; the forces of nature (Einstein)

Reconcile GR with quantum theory

Quantum gravity

Some progress

Black hole thermodynamics

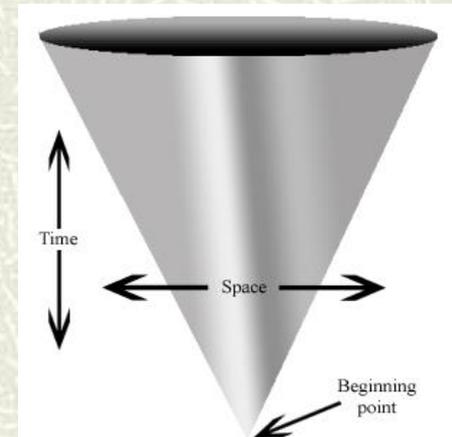
Hawking-Bekenstein radiation

Quantum cosmology

The quantum big bang



Stephen Hawking



A universe from nothing?

Einstein-de Sitter model (1932)

Curvature not a given in dynamic models

Not observed empirically

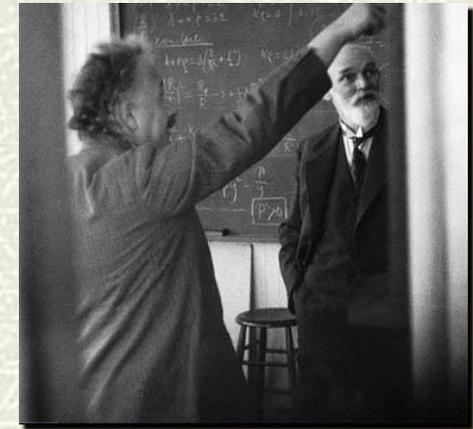
Remove spatial curvature (Occam's razor)

$$ds^2 = -R^2(dx^2 + dy^2 + dz^2) + c^2dt^2$$

$$\frac{3R'^2}{R^2} + \frac{3c^2}{R^2} - \lambda = \kappa c^2 \rho,$$

$$\frac{1}{R^2} \left(\frac{dR}{cdt} \right)^2 = \frac{1}{3} \kappa \rho.$$

$$h^2 = \frac{1}{3} \kappa \rho$$



Simplest Friedman model

Time-varying universe with $\lambda = 0$, $k = 0$

Important hypothetical case: critical universe

Critical density : $\rho = 10^{-28} \text{ g/cm}^3$

Becomes standard model

Despite high density of matter

Despite age problem

Time evolution not considered in paper – see title

PROCEEDINGS
OF THE
NATIONAL ACADEMY OF SCIENCES

Volume 18

March 15, 1932

Number 3

ON THE RELATION BETWEEN THE EXPANSION AND THE
MEAN DENSITY OF THE UNIVERSE

BY A. EINSTEIN AND W. DE SITTER

Communicated by the Mount Wilson Observatory, January 25, 1932

In a recent note in the *Göttinger Nachrichten*, Dr. O. Heckmann has pointed out that the non-static solutions of the field equations of the general theory of relativity with constant density do not necessarily imply a positive curvature of three-dimensional space, but that this curvature may also be negative or zero.

Einstein-de Sitter model revisited

Über das sogenannte kosmologische Problem.

Wenn wir Raum und Zeit von relativistischer Physik absolut nehmen, so hat das folgende Bedeutung. Erstens hat der Raum und Zeit ein bestimmtes, ohne die Bedeutung von einer Realität von dem die Masse. Die Koordinaten hängen auf das gewählte Bezugssystem ab. Zweitens nimmt die Bedeutung der Geometrie und Kosmetik bedingt durch die Relationen zwischen Messungen, welche die Bedeutung von physikalischen Behauptungen haben, die nicht oder falsch sein können. Das Kosmologische Problem ist eine Realität und seine Lösung in der Tragheitsgesetz eingetragene ist das physikalische Problem, was mit dem Raum & Zeit bezeichnet wird, in einem Geometrischen unabhängig von dem Verhalten des übrigen physikalischen Raumes, d. h. der Unabhängigkeit von den Körpern. Der Begriff der Beziehungen zwischen Massenstellen, die alle nur die Masse und Masse zu gewinnen sind, ist unabhängig von der Entstehung und Lösung der Körper unabhängig, ebenso das Kosmologische Problem. In Physik Raum ist gewissermaßen physikalisch zu verstehen oder nicht physikalisch beeinflussbar.

✚ Einstein's cosmology review of 1933

Review of dynamic models from first principles

Culminates in Einstein-de Sitter model

Cosmic constant banished

Possibility of flat geometry

$$2A \frac{d^2A}{dt^2} + \left(\frac{dA}{dt}\right)^2 = 0$$

$$3 \left(\frac{dA}{A}\right)^2 = \kappa \rho c^2.$$

✚ Parameters extracted

Critical density of 10^{-28} g/cm³ : reasonable

Timespan of 10^{10} years: conflict with astrophysics

Attributed to simplifications (incorrect estimate)

$$3h^2 = \kappa \rho c^2 (= 8\pi K \rho)$$

$$A = c(t - t_0)^{\frac{2}{3}}$$

✚ Published in 1933!

French book; small print run

Intended for scientific journal; not submitted

Significant paper



SUR LA STRUCTURE COSMOLOGIQUE DE L'ESPACE ⁽¹⁾

Si nous appelons l'espace et le temps de la physique prérelativiste « absolus », il faut y voir la signification suivante. Tout d'abord l'espace et le temps et, par suite, le système de référence, y figurent dans le même sens comme réalité que, par exemple, la masse. Les coordonnées du système de référence choisi y correspondent immédiatement à des résultats de mesure (2). Les propositions de géométrie et de cinématique signifient pour cette raison des relations entre des mesures ayant la valeur d'affirmations physiques, qui peuvent être vraies ou fausses. Le système d'inertie possède une réalité physique, parce que son choix entre dans la loi d'inertie. En second lieu, cette réalité physique, qui est désignée par les termes espace + temps, est, quant à ses lois, indépendante du comportement des autres réalités physiques, par exemple, des corps.

Albert Einstein



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Albert Einstein

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JPG

Über das sogenannte kosmologische Problem.

by Einstein, Albert (Author)

Date: 1932-09-01

Archival Call Number: 1-115

Document Type: Autograph Draft of Document (ADDf)



DB Info

Kosmologische Betrachtungen zur allgemeinen Relativitätstheorie.

by Einstein, Albert (Author)

Date: 1917-02-08

Archival Call Number: 90-9

Document Type: Printed Document (PD)



DB Info

Die Beantwortung Ihrer Frage, überhaupt kosmologischer Fragen

by Einstein, Albert (Author)

Date: 1929-09-20

Archival Call Number: 25-231

Document Type: Carbon/File Copy of Typed Letter (TLC)



DB Info

Das kosmologische Glied soll überholt sein.

by Hopf, Ludwig (Author)

Date: 1932-06-14

Archival Call Number: 13-306

Document Type: Autograph Letter Signed (ALS)

Albert Einstein

Über das sogenannte...

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Über das sogenannte kosmologische Problem.

Archival Call Number: 1-115

Begin Date: 1932-09-01

End Date: 1932-09-30

Main Author: Einstein, Albert (Author)

Other Persons: Mayer, Walther (Author)

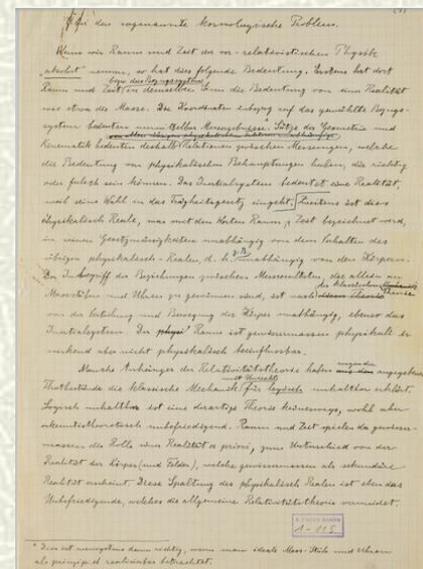
Language: German

Archival Location: Albert Einstein Archives, The Hebrew University of Jerusalem, Israel

Number of Pages: 11.

Document Type Related Items Associated Documents Copyright

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* Dies ist ungenauer Ausdruck, wenn man sich den Text und Thema als geringfügig verändertes betrachtet.

Abandoned model

de Sitter line element

Correct geometry

Simultaneous equations

Error in derivation

Null solution

Einstein's crossroads

Realised problem on revision

Declined to amend model

Evolving models

Less contrived and set $\lambda = 0$

Im Nachfolgenden will ich auf eine Lösung der Gleichung (1) aufmerktsamer machen, welche Hubble's Thatsache gerecht wird, und in welcher die Dichte zeitlich konstant ist. Diese Lösung ist zwar in dem allgemeinen Schema Tolman's enthalten, scheint aber bisher nicht in Betracht gezogen worden zu sein.

1. Ich setze an

$$ds^2 = -e^{\alpha t} (dx_1^2 + dx_2^2 + dx_3^2) + c^2 dt^2 \dots (3)$$

Die Gleichungen (1) liefern

$$-\frac{3}{4} \alpha^2 + \lambda c^2 = 0$$
$$\frac{3}{4} \alpha^2 - \lambda c^2 = \kappa \rho c^2$$

oder

$$\alpha^2 = \frac{\kappa}{3} \rho c^2 \dots (4)$$

Die Dichte ist also konstant und bestimmt die Expansion bis auf das Vorzeichen.

Der Erhaltungssatz bleibt dadurch unvabrt, dass bei Setzung des λ -Gleides der Raum selbst nicht energetisch leer ist; seine Ueltung wird bekanntlich durch die Gleichungen (1) gewährleistet.

A new line of evidence

Gamow and nuclear physics (1940s)

Student of Friedman

How were the chemical elements formed?

Problems with stellar nucleosynthesis

Elements formed in the infant hot universe?

Theory predicts $U = 75\%$ Hydrogen, 25% Helium

Agreement with observation

Victory for big bang model



Georges Gamow



Heavier atoms formed in stars

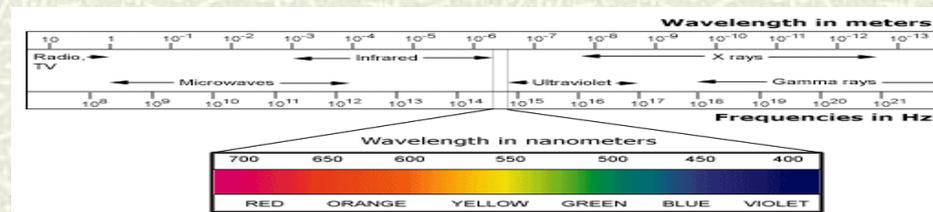
Bonus: cosmic background radiation

- # **Infant universe very hot**
- # **Dominated by radiation**
- # **Radiation still observable today?**
Low temp, microwave frequency
- # **A fossil from the early universe!**
Released when atoms formed (300,000 yr)

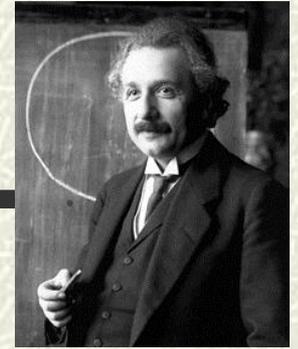
No-one looked



Alpher, Gamow and Herman



Einstein's cosmology: conclusions



Major test for general relativity

Conscious of assumptions of homogeneity, isotropy

Embraces dynamic cosmology

New evidence – new models (JMK)

Timespan of expanding models puzzling

Steady-state universe?



Hubble constant revised

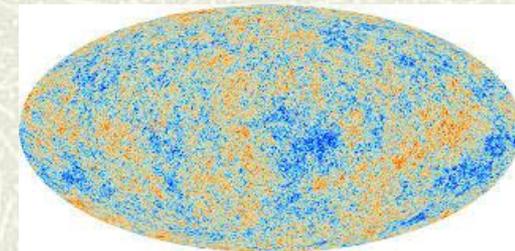
Evolving models (less contrived)

Cosmic constant not necessary

Extraction of parameters compatible with observation

Closed and open models

Timespan problem attributed to simplifying assumptions



Cosmic microwave background
Homogeneous, flat universe

Verdict (1933, 1945): more observational data needed

No mention of origins

Einstein's steady-state model (Jan 31)

Problem with evolving models

“De Sitter and Tolman have already shown that there are solutions to equations (1) that can account for these [Hubbel's] observations. However the difficulty arose that the theory unvaryingly led to a beginning in time about 10^{10} – 10^{11} years ago, which for various reasons seemed unacceptable.”

New solution

“In what follows, I wish to draw attention to a solution to equation (1) that can account for Hubbel's facts, and in which the density is constant over time..

If one considers a physically bounded volume, particles of matter will be continually leaving it. For the density to remain constant, new particles of matter must be continually formed within that volume from space “

Mechanism

“The conservation law is preserved in that, by setting the λ -term, space itself is not empty of energy; its validity is well known to be guaranteed by equations (1).”

Some key quotes (Einstein 1931)

“The cosmological problem is understood to concern the question of the nature of space and the manner of the distribution of matter on a large scale, where the material of the stars and stellar systems is assumed for simplicity to be replaced by a continuous distribution of matter.”

“Now that it has become clear from Hubbel’s results that the extra-galactic nebulae are uniformly distributed throughout space and are in dilatory motion (at least if their systematic redshifts are to be interpreted as Doppler effects), assumption (2) concerning the static nature of space has no longer any justification....”

“Several investigators have attempted to account for the new facts by means of a spherical space whose radius P is variable over time. The first to try this approach, uninfluenced by observations, was A. Friedman,¹ on whose calculations I base the following remarks. ”

“However, the greatest difficulty with the whole approach... is that according to (2 a), the elapsed time since $P = 0$ comes out at only about 10^{10} years. One can seek to escape this difficulty by noting that the inhomogeneity of the distribution of stellar material makes our approximate treatment illusory.”

A useful find

New perspective on steady-state theory (1950s)

Logical idea: not a crank theory

Tolman, Schroedinger, Mimura : considered steady-state universe

Insight into scientific progress

Unsuccessful theories important in the development of science

Links with modern cosmology

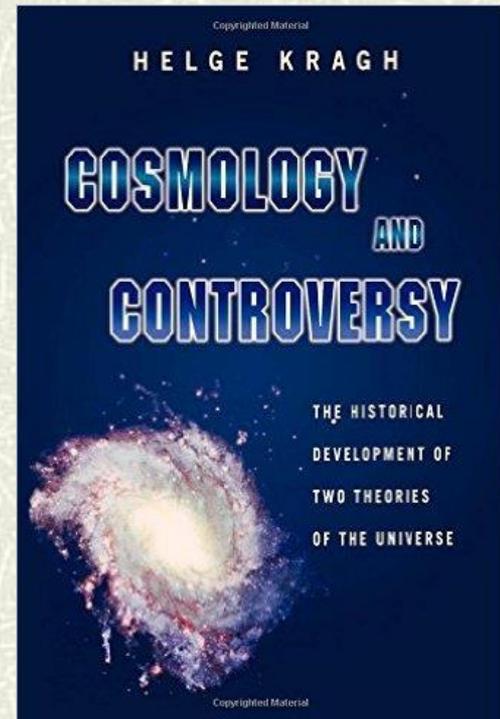
Creation energy and λ : dark energy

de Sitter metric: cosmic inflation

Insight into Einstein's cosmology

Turns to evolving models rather than introduce new term to GFE

Pragmatic approach: F-E model



Einstein's greatest hits (cosmology)



Einstein's model of the Static Universe (1917)

First relativistic model of the cosmos

Einstein's steady-state model (Jan 31)

Natural successor to static model: abandoned

Friedman-Einstein model of the Universe (1931)

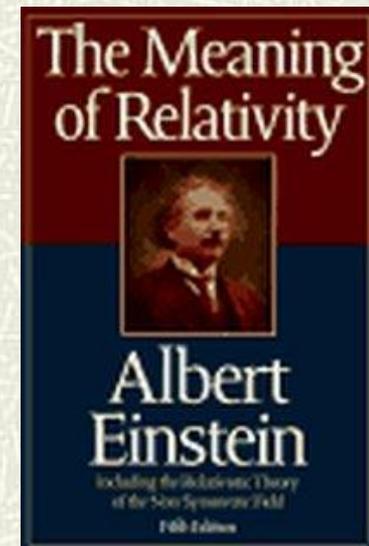
Use of Hubble constant to extract observational parameters

Einstein-de Sitter model of the Universe (1932)

1933 review: 1945 review (Appendix)

Conversations with Gamow, Godel, Straus

No mention of origins



III Astronomy and the Universe

‡ The Great Debate (1900-1925)

Spiral nebulae = galaxies beyond Milky Way?

‡ The Hooker telescope (1917)

Edwin Hubble (1921)

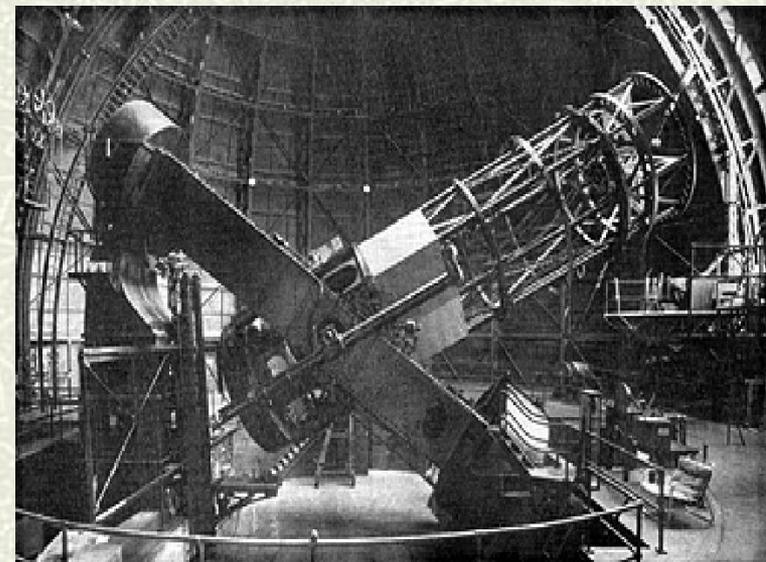
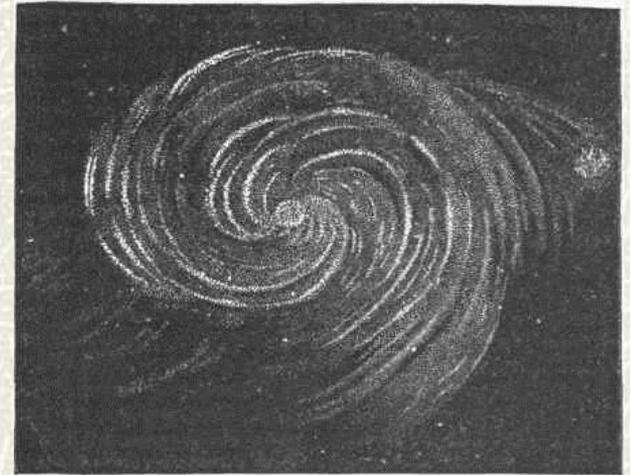
‡ The distances of the nebulae (1925)

Cepheid variables resolved in two nebulae

Leavitt's period-luminosity relation

‡ Spirals far beyond Milky Way

A universe of galaxies



The motion of the nebulae

The redshift of the nebulae

V.M Slipher (Lowell Observatory)

Light from most nebulae redshifted (1915, 1917)

Doppler effect

*Frequency of light depends on
motion of source relative to observer*

Nebulae moving outward?

Galaxies moving outward?

red shift



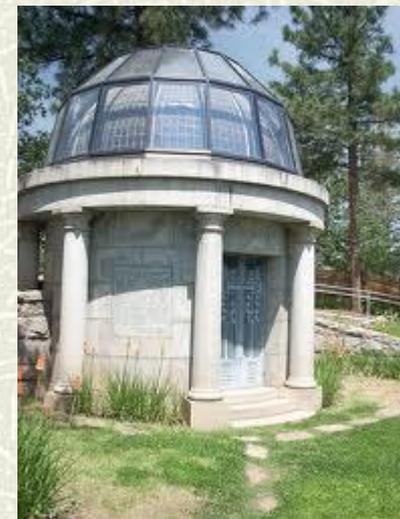
no motion



blue shift



Vesto Slipher



Lowell Observatory

The runaway galaxies (1929)



Edwin Hubble (1889-1953)

A relation between redshift and distance for the galaxies?

Combine 24 distances with redshifts

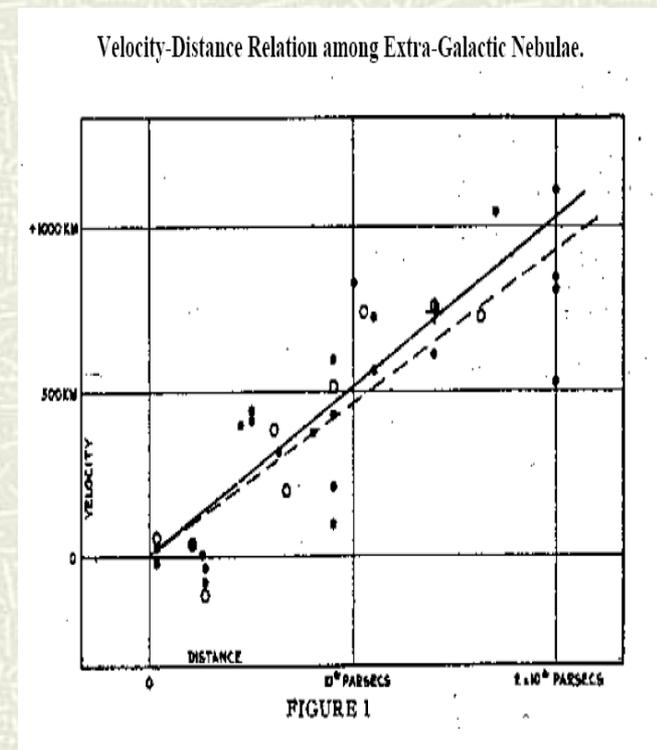
Redshifts from Slipher: not acknowledged

Linear relation: Hubble's law (1929)

$$v = H_0 d \quad \text{with } H = 500 \text{ kms}^{-1} \text{Mpc}^{-1}$$

Landmark result in astronomy

*Far-away galaxies rushing away
at a speed proportional to distance*



Why ?

Lemaître's universe (1927)



Fr Georges Lemaître

Expanding model of the cosmos from GR

Similar to Friedman 1922 model

Starts from static Einstein universe

Recession of nebulae = expansion of space?

Redshifts from Slipher, distances from Hubble

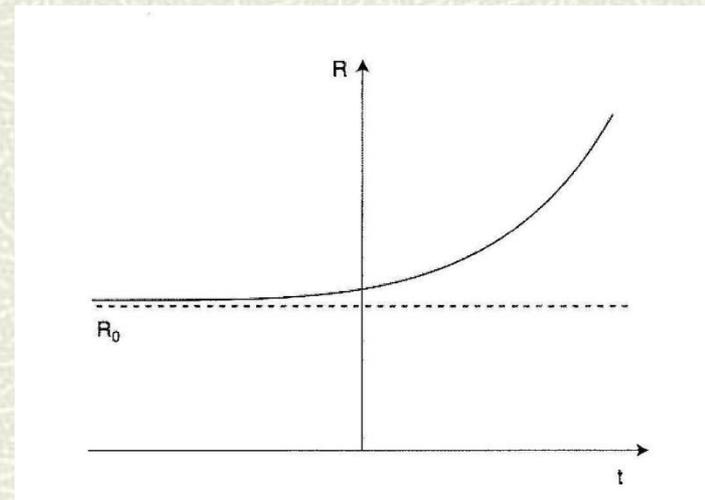
$$H = 585 \text{ kms}^{-1}\text{Mpc}^{-1}$$

Ignored by community

Belgian journal (in French)

Rejected by Einstein: "Votre physique est abominable"

Einstein not up-to-date with astronomy?



The expanding universe (1930)

- **RAS meeting (1930)**

Eddington, de Sitter

If redshifts are velocities, and if effect is non-local

Static cosmic models don't match observations

- **Expanding universe?**

Hubble's law = expansion of space?

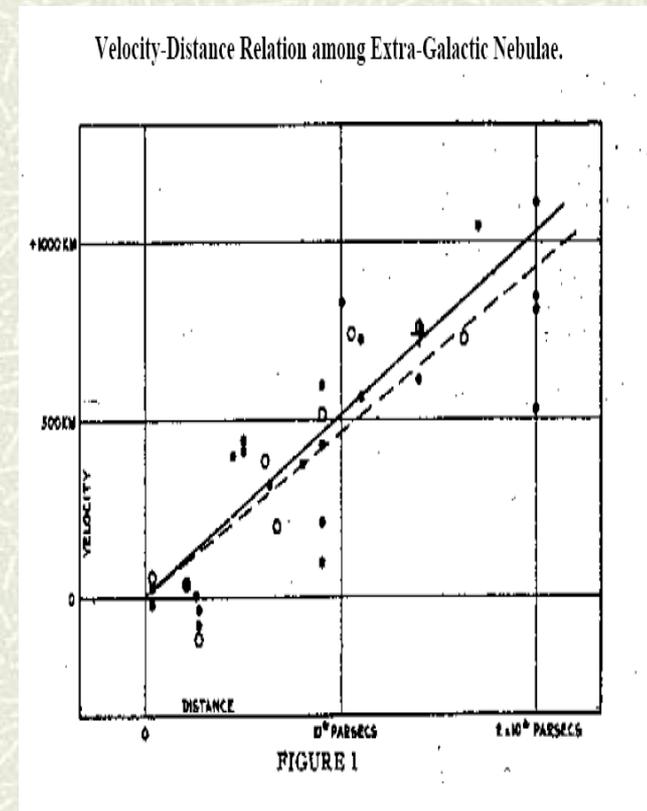
$$H = 500 \text{ kms}^{-1} \text{Mpc}^{-1}$$

- **Friedman-Lemaître model circulated**

Time-varying radius

Time-varying density of matter

Evolving universe



Models of the expanding universe (1930 -)

- **Eddington (1930, 31)**

*On the instability of the Einstein universe
Expansion caused by condensation?*

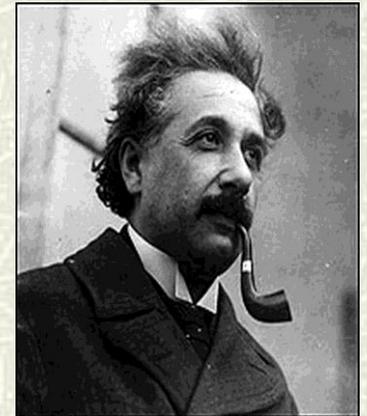
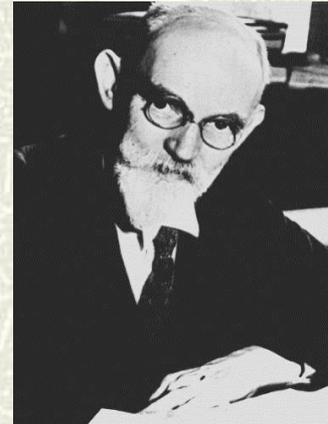


- **Tolman (1930, 31)**

*On the behaviour of non-static models
Expansion caused by annihilation of matter ?*

- **de Sitter (1930, 31)**

*Further remarks on the expanding universe
Expanding universes of every flavour*



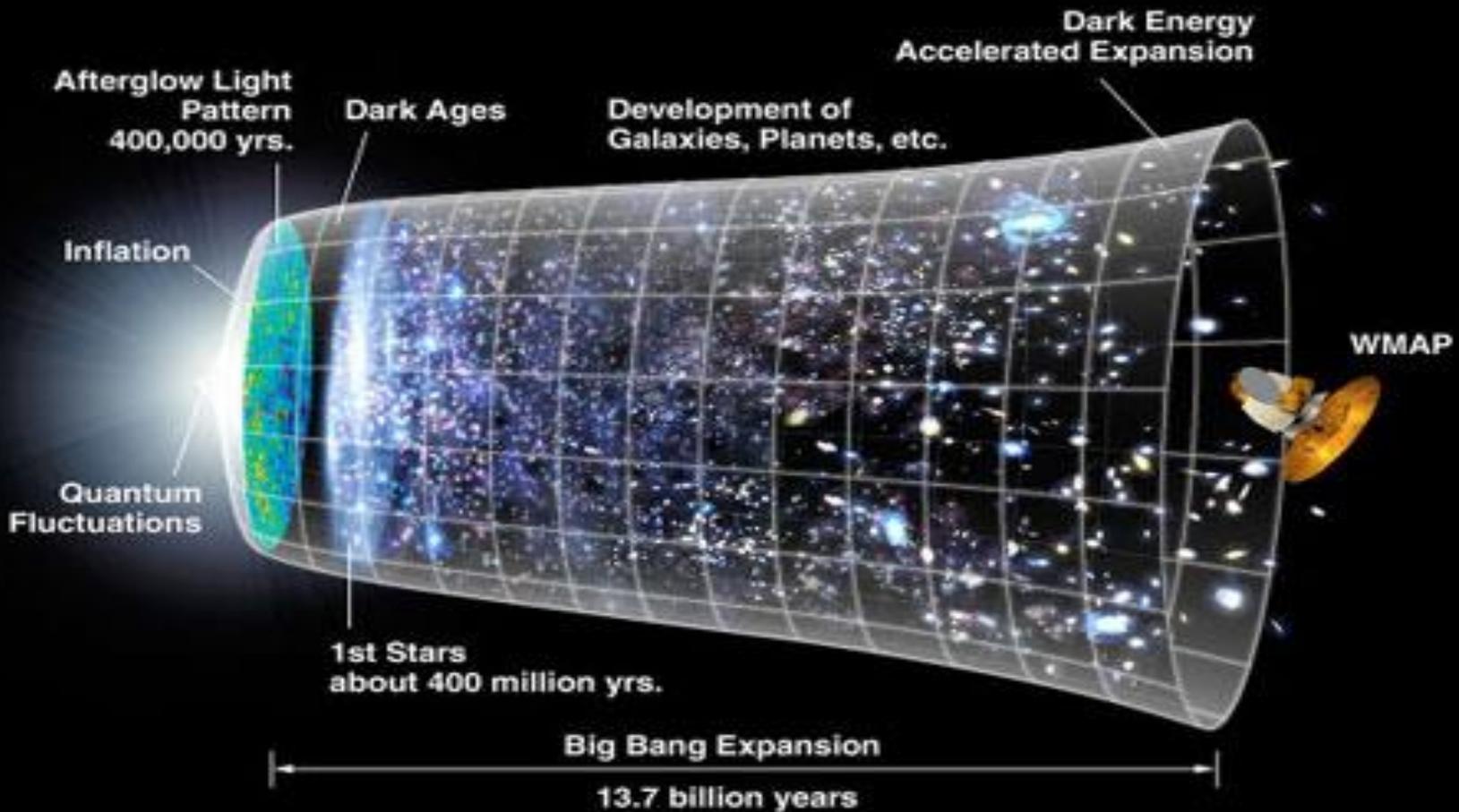
- **Einstein (1931, 32)**

*Friedman-Einstein model $\lambda = 0, k = 1$
Einstein-de Sitter model $\lambda = 0, k = 0$*

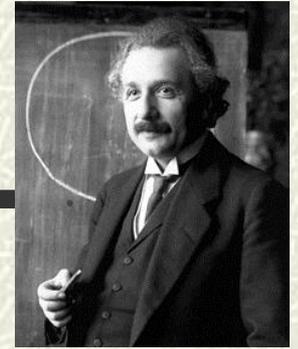
Occam's razor?

*Evolving models
No mention of origins*

The big bang model



Einstein's universe: conclusions



Cosmology = test for general relativity

Introduces λ -term to the field equations

Embraces dynamic cosmology

New evidence – new models

Steady-state vs evolving universe

Evolving models simpler: remove λ -term

The evolving universe

Extract observational parameters

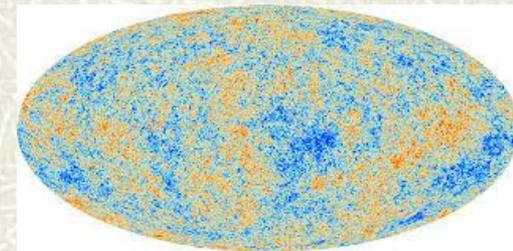
Timespan problem attributed to simplifying assumptions

No discussion of origins

Wary of extrapolations



Hubble constant revised



Cosmic microwave background
Homogeneous, flat universe

Einstein's steady-state model: key quotes

New solution

“In what follows, I wish to draw attention to a solution to equation (1) that can account for Hubbel’s facts, and in which the density is constant over time”

Matter creation

“If one considers a physically bounded volume, particles of matter will be continually leaving it. For the density to remain constant, new particles of matter must be continually formed within that volume from space “

Dark energy

“The conservation law is preserved in that, by setting the λ -term, space itself is not empty of energy; its validity is well known to be guaranteed by equations (1).”

Einstein's steady-state theory: a significant find?

New perspective on steady-state theory (1950s)

Logical possibility: not a crank theory

Insight into scientific progress

Evolution of successful theories

No Kuhnian paradigm shift to 'big bang' model

Slow dawning

Insight into Einstein's philosophy

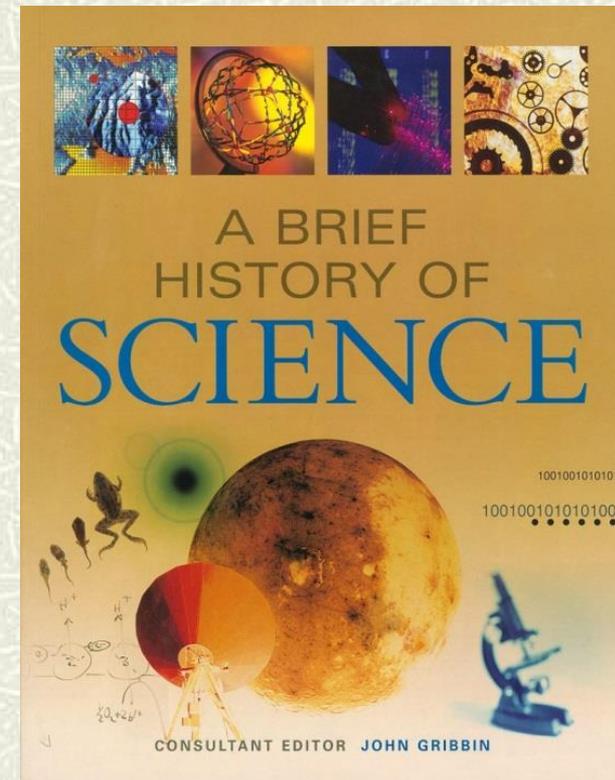
Simple solution?

Discards model rather than introduce new term to GFE

Occam's razor approach

Links with modern cosmology

Dark energy, cosmic inflation



*Paradigm shift or
slow dawning ?*

Explanation for runaway galaxies?

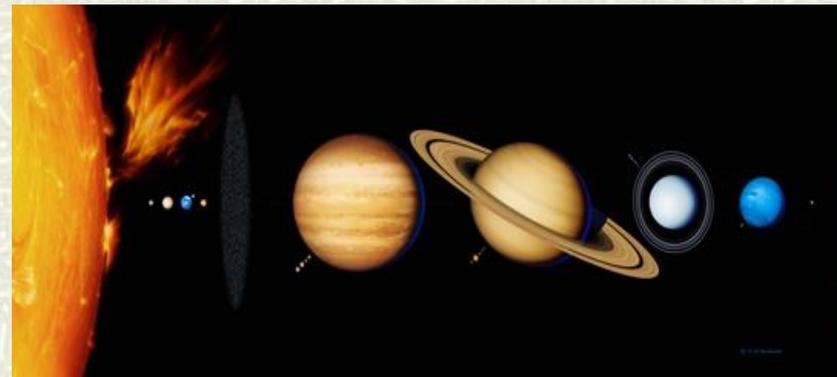
Newton

- Gravity pulls in not out
- Space is fixed
- Time has no beginning

*How can galaxies be receding?
What is pushing out?*



Isaac Newton



Results: publications

■ Einstein's 1931 model

Einstein's cosmic model of 1931 revisited; an analysis and translation of a forgotten model of the universe. O'Raifeartaigh, C. and B. McCann. 2014 *Eur. Phys. J (H)* 39(1):63-85

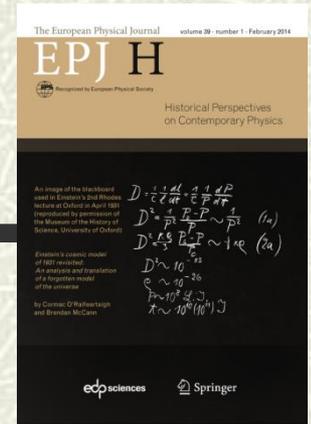
■ Einstein's steady-state manuscript

Einstein's steady-state theory: an abandoned model of the cosmos. O'Raifeartaigh, C., B. McCann, W. Nahm and S. Mitton. 2014 *Eur. Phys. J (H)* 39(3):353-367

■ Einstein-de Sitter model

Einstein's cosmology review of 1933: a new perspective on the Einstein-de Sitter model of the cosmos. O'Raifeartaigh, C., M.O'Keefe, W. Nahm and S. Mitton. 2015. To be published in *Eur. Phys. J (H)*

■ Review paper: conclusions





Edited by
Michael J. Way and Deidre Hunter

Einstein's cosmic model of 1931 revisited: an analysis and translation of a forgotten model of the universe

C. O'Riافةartaigh^a and B. McCann

Department of Computing, Maths and Physics, Waterford Institute of Technology,
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Received 21 September 2013 / Received in final form 20 December 2013
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Abstract. We present an analysis and translation of Einstein's 1931 paper "Zum kosmologischen Problem der allgemeinen Relativitätstheorie" or "On the cosmological problem of the general theory of relativity". In this little-known paper, Einstein proposes a cosmic model in which the universe undergoes an expansion followed by a contraction, quite different to the monotonically expanding Einstein-de Sitter model of 1932. The paper offers many insights into Einstein's cosmology in the light of the first evidence for an expanding universe and we consider his views of issues such as the curvature of space, the cosmological constant, the singularity and the timespan of the expansion. A number of original

An image of the blackboard used in Einstein's 2nd Rhodes lecture at Oxford in April 1931 (reproduced by permission of the Museum of the History of Science, University of Oxford)

$$D = \frac{1}{c} \frac{dL}{dt} = \frac{1}{c} \frac{dP}{d\tau}$$

$$D^2 = \frac{1}{P^2} \frac{P_0 - P}{P} \sim \frac{1}{P^2} \quad (1a)$$

$$D^2 = \frac{K_0}{3} \frac{P_0 - P}{P} \sim \frac{1}{4} \kappa P \quad (2a)$$

$$D^2 \sim 10^{-53}$$

$$c \sim 10^{10} \text{ s}^{-1}$$

$$P \sim 10^8 \text{ s}$$

$$\tau \sim 10^{10} (10^{11}) \text{ s}$$

Einstein's cosmic model of 1931 revisited:
An analysis and translation of a forgotten model of the universe

by Cormac O'Riافةartaigh and Brendan McCann

Einstein's steady-state theory: an abandoned model of the cosmos

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Published online (Inserted Later)
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Abstract. We present a translation and analysis of an unpublished manuscript by Albert Einstein in which he attempted to construct a "steady-state" model of the universe. The manuscript, which appears to have been written in early 1931, demonstrates that Einstein once considered a cosmic model in which the mean density of matter in an expanding universe is maintained constant by the continuous formation of matter from empty space. This model is very different to previously

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Comments: Accepted for publication in the European Physical Journal (H). Includes an English translation of a little-known review of cosmology written by Albert Einstein in 1933. 20 pages, 4 figures

Subjects: History and Philosophy of Physics (physics.hist-ph); Cosmology and Nongalactic Astrophysics (astro-ph.CO)

Cite as: arXiv:1503.08029 [physics.hist-ph] (or arXiv:1503.08029v1 [physics.hist-ph] for this version)

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Comments: 20 pages, 2 figures. To be published in the book The Philosophy of Cosmology: Foundations and Perspectives (Cambridge University Press)

Subjects: History and Philosophy of Physics (physics.hist-ph); Cosmology and Nongalactic Astrophysics (astro-ph.CO)

Cite as: arXiv:1504.02873 [physics.hist-ph] (or arXiv:1504.02873v1 [physics.hist-ph] for this version)

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[v1] Sat, 11 Apr 2015 13:39:48 GMT (653K)

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Die Gleichungen (1) liefern

$$-\frac{3}{4} \alpha^2 + \lambda c^2 = 0$$

$$\frac{3}{4} \alpha^2 - \lambda c^2 = \kappa \rho c^2$$

oder

$$\alpha^2 = \frac{\kappa \rho c^2}{3} \quad \dots \dots (4)$$

Die Dichte ist also konstant und bestimmt die Expansion bis auf das Vorzeichen.

Taking $T_{44} = \rho c^2$ (all other components zero) in the *time* component of equation (1) we obtain $\left(R_{44} - \frac{1}{2} g_{44} R\right) - \lambda g_{44} = \kappa \rho c^2$.

This gives on analysis $-\frac{3\alpha^2}{4} + \frac{3\alpha^2}{2} - \lambda c^2 = \kappa \rho c^2$ the second of Einstein's simultaneous equations.

From the *spatial* component of equation (1), we obtain $\left(R_{ii} - \frac{1}{2} g_{ii} R\right) - \lambda g_{ii} = 0$.

This gives on analysis $\frac{3\alpha^2}{4} - \frac{3\alpha^2}{2} + \lambda c^2 = 0$ for the first of the simultaneous equations.

It is plausible that Einstein made a sign error here, initially getting $\frac{3\alpha^2}{4} + \frac{3\alpha^2}{2} + \lambda c^2 = 0$ for this equation. (W. Nahm)

Einstein's steady-state model and cosmology today

Accelerated expansion (1998)

Supernova measurements

Dark energy – positive cosmological constant



Einstein's dark energy

“The conservation law is preserved in that, by setting the λ -term, space itself is not empty of energy; its validity is well known to be guaranteed by equations (1).”

Anticipates positive cosmological constant

De Sitter line element

$$ds^2 = - e^{at} (dx_1^2 + dx_2^2 + dx_3^2) + c^2 dt^2 \dots$$

Necessary for all steady-state models

Identical to inflationary models (different time-frame)

Some key quotes (Einstein 1917)

“In a consistent theory of relativity, there can be no inertia relative to “space”, but only an inertia of masses relative to one another”

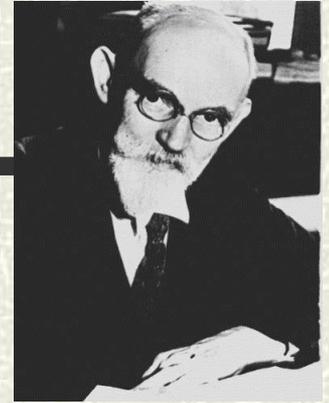
“I have not succeeded in formulating boundary conditions for spatial infinity. Nevertheless, there is still a way out...for if it were possible to regard the universe as a continuum which is finite (closed) with respect to its spatial dimensions, we should have no need at all of any such boundary conditions”

“The most important fact that we draw from experience as to the distribution of matter is that the relative velocities of the stars are very small compared with the velocity of light..... There is a system of reference relative to which matter may be looked upon as being permanently at rest ”

“However, the system of equations ..allows a readily suggested extension which is compatible with the relativity postulate... For on the left hand side of the field equation...we may add the fundamental tensor $g_{\mu\nu}$, multiplied by a universal constant , $-\lambda$, at present unknown, without destroying the general covariance ”

Schroedinger's comment (1918): Einstein's response (1918)

de Sitter's universe



Alternative solution of the GFE

A universe empty of matter (1917)

Solution B

Cosmic constant proportional to curvature of space

$$\lambda = 3/R$$

Disliked by Einstein

Conflict with Mach's principle

Problems with singularities? (1918)

The de Sitter confusion

Static or non-static - a matter of co-ordinates?

Weyl, Lanczos, Klein, Lemaître

[p. 270] 5. "Critical Comment on a Solution of the Gravitational Field Equations Given by Mr. De Sitter"

[Einstein 1918c]

SUBMITTED 7 March 1918

PUBLISHED 21 March 1918

IN: *Königlich Preußische Akademie der Wissenschaften (Berlin). Sitzungsberichte* (1918): 270–272.

[1] Herr De Sitter, to whom we owe deeply probing investigations into the field of the general theory of relativity, has recently given a solution for the equations of gravitation which, in his opinion, could possibly represent the metric structure of the universe. However, it appears to me that one can raise a grave argument against the admissibility of this solution, which shall be presented in the following.

The De Sitter solution of the field equations

$$G_{\mu\nu} - \lambda g_{\mu\nu} = -\kappa T_{\mu\nu} + \frac{1}{2}g_{\mu\nu}\kappa T \quad (1)$$

is

Prediction of redshifts – Slipher effect?

An abandoned model

✦ Correct geometry

de Sitter metric

✦ Simultaneous equations

Eliminate λ

Relation between α^2 and ρ

✦ Einstein's crossroads

Null solution on revision

Tolman? (Nussbaumer 2014)

Declined to amend GFE

✦ Evolving models

Less contrived: set $\lambda = 0$

Im Nachfolgenden will ich auf eine Lösung der Gleichung (1) aufmerktsamer machen, welche Hubble's Thatsache gerecht wird, und in welcher die Dichte zeitlich konstant ist. Diese Lösung ist zwar in dem allgemeinen Schema Tolman's enthalten, scheint aber bisher nicht in Betracht gezogen worden zu sein.

1. Ich setze an

$$ds^2 = -e^{\alpha t} (dx_1^2 + dx_2^2 + dx_3^2) + c^2 dt^2 \dots (3)$$

Die Gleichungen (1) liefern

$$-\frac{3}{4} \alpha^2 + \lambda c^2 = 0 \qquad 9\alpha^2 / 4 + \lambda c^2 = 0$$

$$\frac{3}{4} \alpha^2 - \lambda c^2 = \kappa \rho c^2 \qquad 3\alpha^2 / 4 - \lambda c^2 = \kappa \rho c^2$$

oder

$$\alpha^2 = \frac{\kappa}{3} \rho c^2 \dots (4) \qquad \alpha^2 = \frac{\kappa c^2}{3} \rho$$

Die Dichte ist also konstant und bestimmt die Expansion bis auf das Vorzeichen.

Der Erhaltungssatz bleibt dadurch unvabert, dass bei Setzung des λ -Gledes der Raum selbst nicht energetisch leer ist; seine Ubertung wird bekanntlich durch die Gleichungen (1) gewahrleistet.

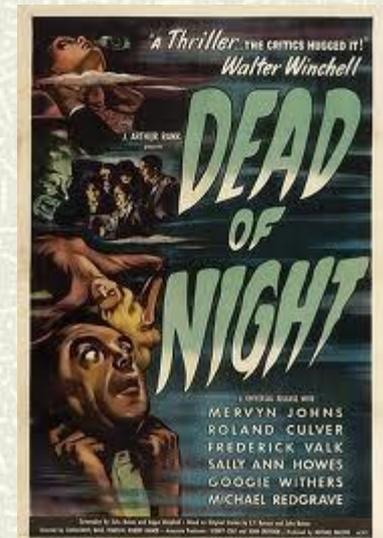
Steady-state universe (1948)

- # Alternative to big bang (*Fred Hoyle*)
- # Expanding universe

BUT

- # Continuous creation of matter?
- # Unchanging universe
- # No beginning, no age problem
- # No assumptions about early epochs

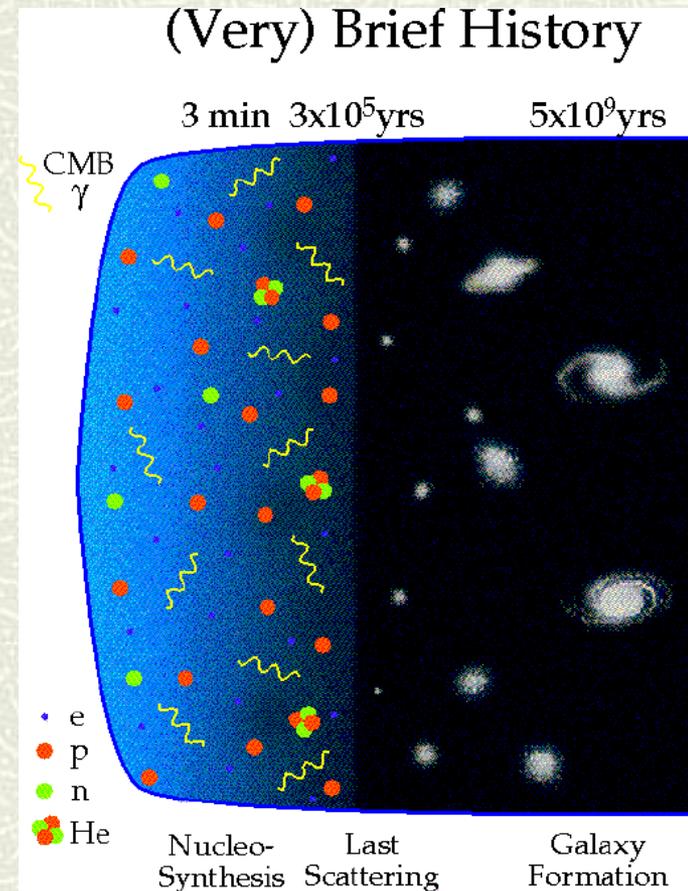
Very little matter needed



The big bang – evidence

1. The expansion of the U
2. The abundance of H and He
3. The distribution of the galaxies
4. The cosmic microwave background

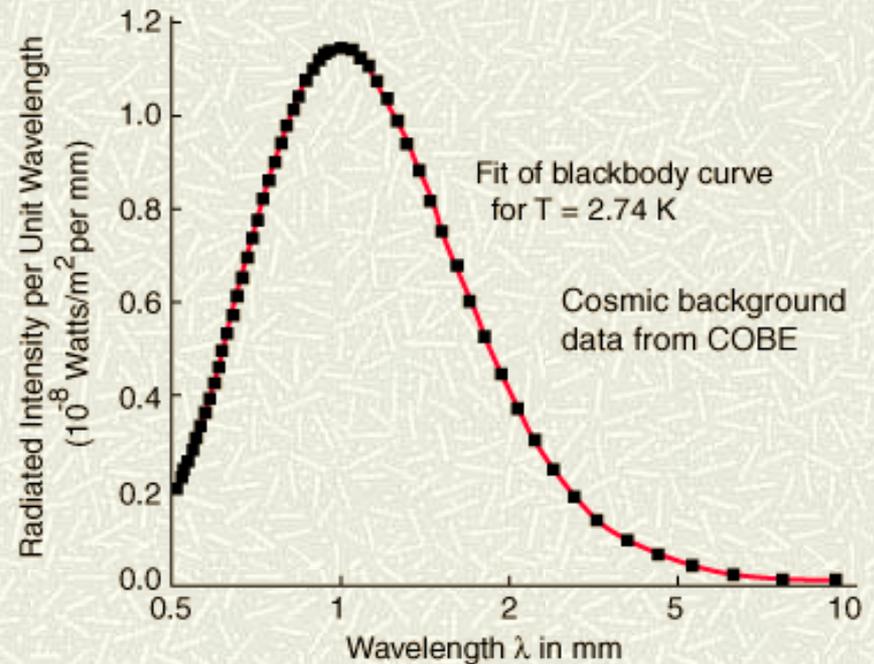
How did it start?



COBE measurements of CMB

- Expected temperature
- Expected frequency
- Perfect blackbody spectrum

- *Radiation very uniform*
- *Variation of 1 in 10^5*
- *Seeds of galaxies ?*



Nobel Prize

COBE (1992)

3. Einstein's steady-state model

Unpublished manuscript

Archived as draft of F-E model (1931)

Similar title, opening to F-E model

Something different

Cosmological constant

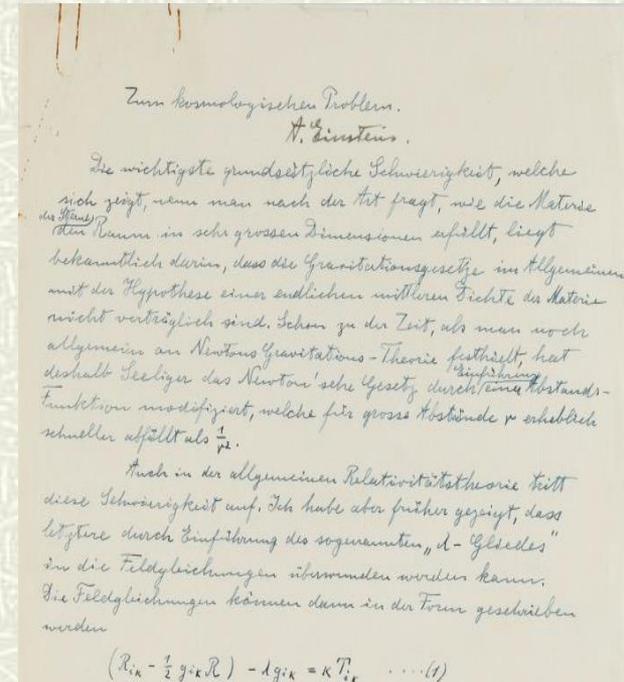
"The density is thus constant and determines the expansion"

Steady-state model of the Expanding Universe

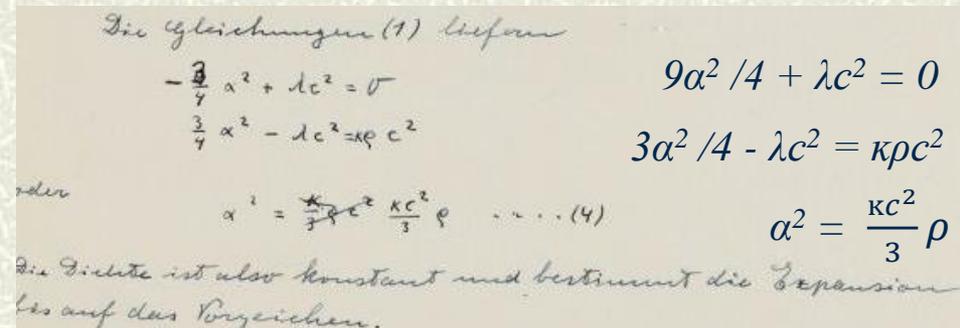
Anticipates Hoyle solution

Written in early 1931

Fatal flaw: abandoned



$$(R_{ik} - \frac{1}{2} g_{ik} R) - \lambda g_{ik} = \kappa T_{ik} \dots (1)$$



NATURE | NEWS   

Einstein's lost theory uncovered

Physicist explored the idea of a steady-state Universe in 1931.

Daive Castelvechi

24 February 2014

Physics » Nature   Email  Print

Einstein's Lost Theory Uncovered

The famous physicist explored the idea of a steady-state universe in 1931

nature

Feb 25, 2014 | By **Daive Castelvechi** and Nature magazine

A manuscript that lay unnoticed by scientists for decades has revealed that **Albert Einstein** once dabbled with an



New Discovery Reveals Einstein Tried To Devise A Steady State Model Of The Universe

  2 comments, 2 called-out [+ Comment Now](#) [+ Follow Comments](#)

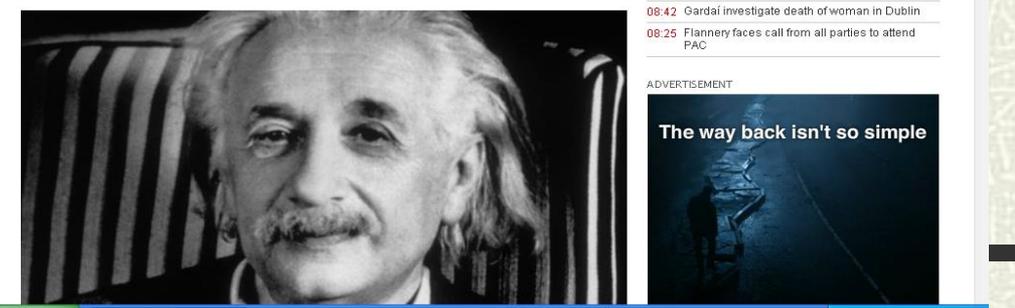
WIT researchers discover 'lost' Einstein model of universe

Scientists uncovered misfiled papers while searching Jerusalem university's online archive

Almost 20 years before the late Fred Hoyle and his colleagues devised the [Steady State Theory](#), Albert Einstein toyed with a similar idea: that the universe was eternal, expanding outward with a consistent input of spontaneously generating matter.

- Latest Ireland »**
- 12:26 Quinn confirms Flannery approached hm with Rehab concerns
 - 09:07 Man in his twenties stabbed in north Dublin
 - 09:05 Family hope public appeal will help daughter beat cancer
 - 08:42 Gardaí investigate death of woman in Dublin
 - 08:25 Flannery faces call from all parties to attend PAC

An Irish physicist came across the paper last year and could hardly believe. According to this week's article in [Nature](#),



model of the universe very different to today's [Big Bang](#) Theory.

The manuscript, which hadn't been referred to by scientists for decades,

 **TheJournal.ie**  Like You like this.

 **Adbank** The straight talking savings bank

2. Einstein-de Sitter model (1932)

Remove spatial curvature

Curvature not a given in dynamic models (Heckmann)

Not observed empirically (Occam's razor)

$$ds^2 = -R^2(dx^2 + dy^2 + dz^2) + c^2dt^2$$

Simplest Friedman model

Time-varying universe with $\lambda = 0$, $k = 0$, $p = 0$

Estimate of density : $\rho = 10^{-28}$ g/cm³

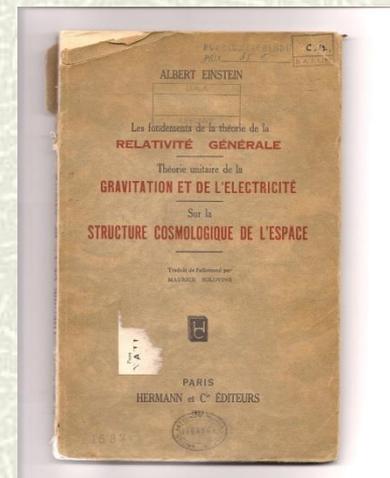
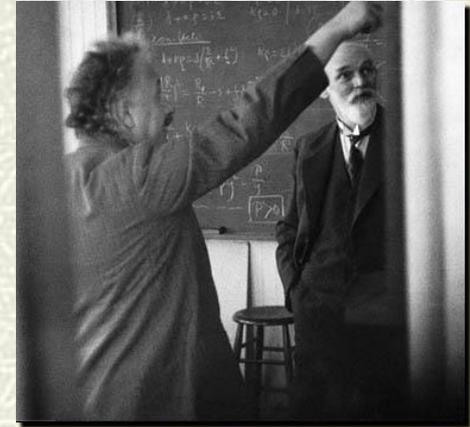
$$\frac{1}{R^2} \left(\frac{dR}{cdt} \right)^2 = \frac{1}{3} \kappa \rho.$$

Becomes standard model

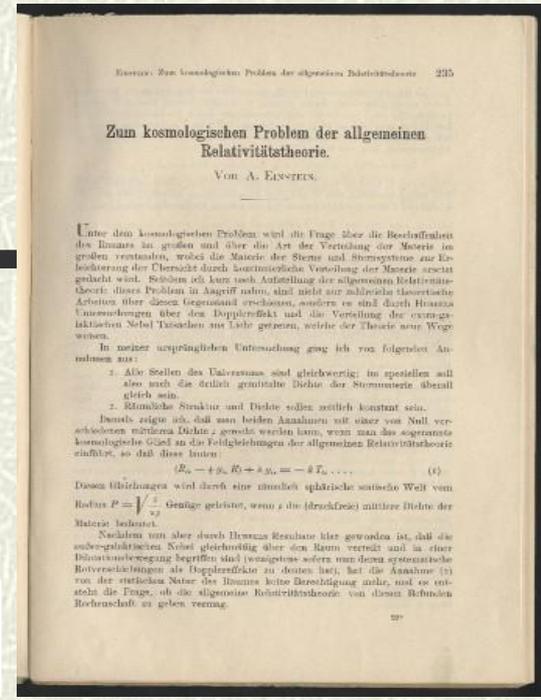
Despite high density of matter, age problem

Time evolution not considered

Longer version with time evolution (Einstein 1933)



Einstein's 1931 model (F-E)



✦ Einstein's first expanding model

Rarely cited (not translated)

$$\frac{3P'^2}{P^2} + \frac{3c^2}{P^2} - \lambda = \kappa c^2 \rho.$$

✦ Adopts Friedman 1922 model

Instability of static solution

Hubble's observations

$$\frac{P'^2}{P^2} + \frac{2P''}{P} + \frac{c^2}{P^2} - \lambda = 0$$

$$\left(\frac{dP}{dt}\right)^2 = c^2 \frac{P_0 - P}{P}$$

✦ Sets cosmic constant to zero

Redundant

$$D^2 = \frac{1}{P^2} \frac{P_0 - P}{P}$$

$$P \sim \frac{1}{D}$$

$$D = \frac{1}{P} \frac{dP}{dt} \cdot \frac{1}{c}$$

✦ Extraction of cosmic parameters

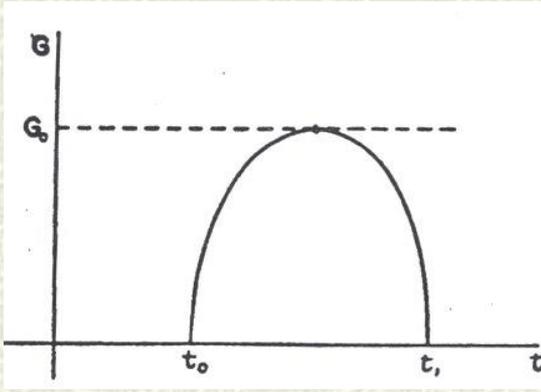
$$P \sim 10^8 \text{ lyr} : \rho \sim 10^{-26} \text{ g/cm}^3$$

t ~ 10¹⁰ yr : conflict with astrophysics

$$D^2 = \frac{1}{3} \kappa \rho \frac{P_0 - P}{P}$$

$$D^2 \sim \kappa \rho$$

Attributed to simplifying assumptions (homogeneity)



Einstein's 1931 model revisited

First translation into English

O'Raifeartaigh and McCann 2014

$$D = \frac{1}{P} \frac{dP}{dt} \cdot \frac{1}{c}$$

$$D^2 = \frac{1}{P^2} \frac{P_0 - P}{P}$$

$$P \sim \frac{1}{D}$$

Not a cyclic model

"Model fails at $P = 0$ "

Contrary to what is usually stated

$$D^2 \sim \kappa \rho$$

*Oxford lecture
(May 1931)*

Anomalies in calculations of radius and density

Einstein: $P \sim 10^8$ yr, $\rho \sim 10^{-26}$ g/cm³, $t \sim 10^{10}$ yr

We get: $P \sim 10^9$ yr, $\rho \sim 10^{-28}$ g/cm³, $t \sim 10^9$ yr

Source of error?

Oxford blackboard: $D^2 \sim 10^{-53}$ cm⁻² should be 10^{-55} cm⁻²

Time miscalculation $t \sim 10^{10}$ yr (should be 10^9 yr)

Non-trivial error: misses conflict with radioactivity

$$D = \frac{1}{c} \frac{1}{P} \frac{dP}{dt}$$

$$D^2 = \frac{1}{P^2} \frac{P_0 - P}{P} \sim \frac{1}{P^2} \quad (1a)$$

$$D^2 = \frac{\kappa g}{3} \frac{P_0 - P}{P} \sim \frac{\kappa g}{3} \quad (2a)$$

$$D^2 \sim 10^{-53}$$

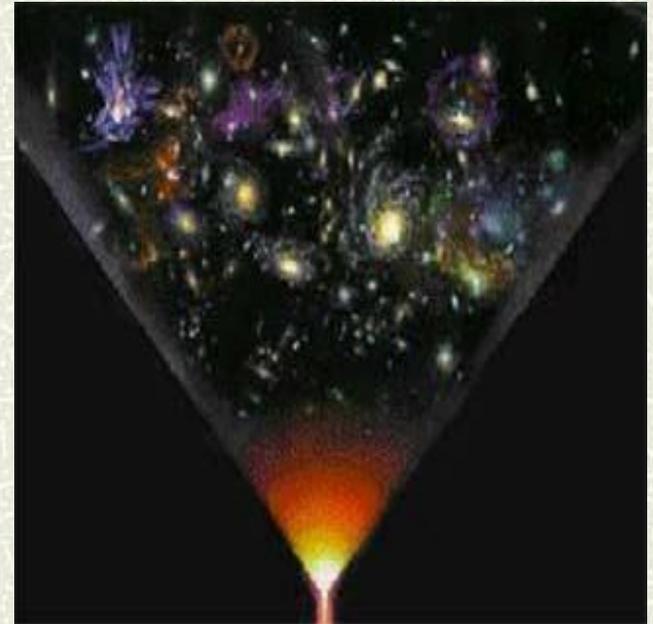
$$\rho \sim 10^{-26}$$

$$P \sim 10^8 \text{ yr}$$

$$t \sim 10^{10} (10^9) \text{ yr}$$

IV The 'big bang' model (1931)

- # Infant U concentrated in tiny volume
- # Extremely dense, hot
- # Expanding and cooling ever since



Where do the laws of physics come from?

Wrong age (Hubble constant)

Singularity problem
 ∞ density, ∞ temp at $t = 0$?